Deep Learning Neural Networks Trained with MODIS Satellite-Derived Predictors for Long-Term Global Solar Radiation Prediction

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Abstract: Solar energy predictive models designed to emulate the long-term (e.g., monthly) global solar radiation (*GSR*) trained with satellite-derived predictors can be employed as decision tenets in the exploration, installation and management of solar energy production systems in remote and inaccessible solar-powered sites. In spite of a plethora of models designed for *GSR* prediction, deep learning, representing a state-of-the-art intelligent tool, remains an attractive approach for renewable energy exploration, monitoring and forecasting. In this paper, algorithms based on deep belief networks and deep neural networks are designed to predict long-term *GSR*. Deep learning algorithms trained with publicly-accessible Moderate Resolution Imaging Spectroradiometer (MODIS) satellite data are tested in Australia's solar cities to predict the monthly *GSR:* single hidden layer and ensemble models. The monthly-scale MODIS-derived predictors (2003–2018) are adopted, with 15 diverse feature selection approaches including a Gaussian Emulation Machine for sensitivity analysis used to select optimal MODIS-predictor variables to simulate *GSR* against ground-truth values. Several statistical score metrics are adopted to comprehensively verify surface *GSR* simulations to ascertain the practicality of deep belief and deep neural networks. In the testing phase, deep learning models generate significantly lower absolute percentage bias (≤3%) and high Kling–Gupta efficiency (≥97.5%) values compared to the single hidden layer and ensemble model. This study ascertains that the optimal MODIS input variables employed in *GSR* prediction for solar energy applications can be relatively different for diverse sites, advocating a need for feature selection prior to the modelling of *GSR*. The proposed deep learning approach can be adopted to identify solar energy potential proactively in locations where it is impossible to install an environmental monitoring data acquisition instrument. Hence, MODIS and other related satellite-derived predictors can be incorporated for solar energy prediction as a strategy for long-term renewable energy exploration.

Keywords: global solar radiation; energy security; deep learning; deep belief network; deep neural network; solar cities in Australia

1. Introduction and Background

Due to the decreasing trends of feed-in tariffs (a premium rate paid for electricity fed back into the electricity grid from a designated renewable electricity generation source) for solar-generated electricity in many countries (including Australia), there has been an accelerated interest and need for versatile energy management systems (EMS) for end-users to increase the generation of

electricity and the capacity for power transmission from various regions, both remote and metropolitan, to meet the rising consumer energy demands [1]. EMS are able to monitor, control and optimize the transmission and use of solar and conventional energies [2]. However, the prediction error on the power output from a solar power system can cause a negative effect on the system profitability. Considering this, an accurate predictive tool for solar can help reduce the uncertainty of power generation (solar photovoltaic), as well as to increase the conversion efficiency (solar thermal) in the future. Such tools can be used to explore and evaluate the sustainability of long-term solar-powered energy installations in all regions, irrespective of their location.

The magnitude of power generated by a solar photovoltaic (PV) system and conversion efficiency (solar air heater, solar water heater, solar concentrator) is largely a function of the global solar radiation (*GSR*) [3]. However, stochastic components of solar energy variability depend on the cloud coverage characteristics, as well as factors including the aerosol, dust particles, smoke and airborne pollutants that are largely difficult to measure on an ongoing basis and, therefore, must be derived from remotely-sensed data products. In addition to this, the intermittency and randomness in atmospheric variables and a lack of data for remote or regional sites make the prediction of long-term availability of *GSR* to support a future solar PV, as well as a solar thermal system quite challenging. Although *GSR* is one of the most commonly-monitored meteorological variables, measurement stations remain sparse, particularly in the Southern Hemisphere [4]. Even if a measurement station has been set up, the measured data can be unreliable and questionable, due to a lack of regular maintenance and issues with calibration of the instruments in a regional or remote location [5]. To surmount these issues, the opportunity to adopt satellite-derived predictors to estimate long-term *GSR* presents an alternative and viable avenue for future exploration of solar energy.

To explore solar energy potentials, many techniques have been developed to predict *GSR*, which can be largely categorized as follows: (I) empirical models with simple mathematical equations: linear, quadratic and polynomial equations to emulate the links between *GSR* and its related meteorological variables; (II) remote sensing retrieval that is based on images from satellites used to predict *GSR* [6]; (III) soft computing or data-driven models that apply artificial intelligence techniques to model the erratic behaviour of *GSR* received on the Earth's surface. The requirement for any predictive model for solar applications is that it must be an appropriately representative model developed, calibrated and validated to extract intrinsic features related to *GSR* prediction.

Data-driven models are becoming increasingly promising tools for electrical power [7,8] and solar radiation prediction [9–14]. Single hidden layer (SHL) neural networks, using the artificial neural network (ANN) as a black-box tool, have been designed for both short-term [15] and long-term prediction of *GSR* [13,16]. A recent study in Australia designed an ANN model at four locations using the European Centre for Medium Range Weather Forecasting (ECMWF) reanalysis data as an input [10]. In spite of the acceptable performance in this study, inputs were selected from a limited set of meteorological variables from weather stations (*e.g.*, latitude, rainfall, sunshine duration, humidity, temperature) and, therefore, did not consider additional predictors, such as those available in satellite data repositories that could possibly influence *GSR*.

To address the potential problems associated with the inadequacy of data, the opportunity to use satellite products from the National Aeronautics and Space Administration (NASA), Goddard Online Interactive Visualization and Analysis Infrastructure (GIOVANNI) repository is an alternative avenue to generate *GSR* forecasts, particularly feeding the model with important variables such as land surface temperature, cloud-free days, aerosol optical depth and cloud temperature that are highly likely to moderate the amount of solar radiation received at the Earth's surface. In fact, recent studies have utilized land surface temperature with other satellite-derived variables to model long-term *GSR* in regional Queensland and over the Australian sub-continent [9,13], although none have used a sophisticated method (*e.g.*, deep learning algorithms).

Recently, to address potential limitations of ANNs, particularly arising from the algorithm being a single hidden layer neuronal system, a number of newer neural network techniques such as deep learning (DL) have also been implemented [17] and shown to generate a superior accuracy

compared to a single hidden layer model. Deep learning is designed to use a neural network structure similar to the ANN to represent inputs and target data. These models use multiple feature extraction layers and learn the complex relationships within the data more efficiently. These DL methods have been widely implemented in medical imaging, speech recognition and natural language processing, autonomous driving and computer vision. However, there have been only a few prior studies that have employed a DL model for *GSR* prediction, especially using satellite-derived predictor datasets.

To address the limitations of single hidden layer neuronal models, this paper adopts the deep neural network (DNN) and deep belief network (DBN), the two fundamental categories of DL algorithms, coupled with satellite-derived data to predict long-term *GSR,* where monthly averaged daily values are modelled for solar cities in Australia. These solar cities have previously been established as potential future sites for solar energy projects that have a low cloud cover and limited aerosol concentrations and are thus well suited for solar energy. To provide a sound context for developing DNN and DBN models to predict the *GSR*, the merits of DL models include the capability to extract much deeper and naturally inherent data features within a predictor-target matrix, mainly to provide more accurate predictions [18]. For example, a DNN approach is able to boost the predictive power of the ANN model by deepening and replicating its hidden layers and also leveraging its internal structures to model the *GSR* accurately. Moreover, a DBN model [19] is able to avoid the problem of overfitting and also avoiding the learning mode being halted when a local optima emerges in a feature space. The merits of deep learning models can, therefore, help address the unavoidable drawbacks of conventional approaches, *e.g.*, an ANN model [20].

Many studies are currently using deep learning for time series forecasts [21–24]. Some results reveal a DBN model's superiority over a linear autoregressive and a conventional back-propagation neural network (*i.e.*, ANN) model. Furthermore, literature on *GSR* prediction problems using deep learning approaches has rather been limited to short-term forecast horizons (*i.e.*, minutes and hours), and these studies have used deep learning based on long short term memory network or convolutional neural networks. However, a longer forecast horizon (*i.e.*, weekly or monthly model) can be useful for exploring the long-term prospects of solar energy [25], leading to better policy, implementation of new solar powered sites and expansion of solar energy systems in remote and regional locations where solar radiation may be in abundance [9,10,13].

A literature review, particularly the related review articles [26–28], shows that the current literature is relatively scarce, and even non-existent, in terms of prior studies conducted to predict monthly *GSR* using deep learning approaches. From a practical point of view, the future planning for an electricity grid certainly requires the prediction of solar radiation a few months ahead of time [29]; therefore, a monthly predictive model is particularly desirable. That model can be useful for agricultural crop growth [30], production of algal-derived biofuels [31], and key decisions made for many applications, where the estimation of long-term solar radiation may be required.

The aims of this study are as follows: (1) to design and implement a deep learning (DL) approach using deep belief network (DBN) and deep neural network (DNN) algorithms and to evaluate its relative success in estimating the long-term daily average monthly *GSR* using remotely-sensed MODIS-derived products as the DL model's input variables. Here, we consider the application study site as Australia's solar cities, namely: Blacktown [33.77°S, 150.90°E], Adelaide [34.92°S, 138.59°E], Townsville [19.25°S, 146.81°E] and Central Victoria [36.74°S, 144.28°E], all four of which are situated in the dry sub-tropic region and are relatively enriched with solar exposure. The next aims of the study are: (2) to apply wrapper and filter-based feature selection techniques on the MODIS satellite data in order to select the optimum predictor variables for these prescribed DL models; (3) to adopt the Gaussian Emulation Machine approach to perform a sensitivity analysis of MODIS variables to deduce their relative influence on *GSR* prediction; (4) to benchmark the deep learning models (*i.e.*, DBN and DNN) with a multitude of competing data-driven approaches, namely: single hidden layer (ANN), and ensemble models (random forest regression (RF), extreme gradient boosting regression (XGBR), Gradient Boosting Machine (GBM), and decision tree (DT).

By testing the developed models over Australia's solar cities, this paper aims to provide valuable contributions to exploring the utility of a deep learning approach of improving other previous studies (*e.g.*, [10–13]) where single hidden layer neuronal systems have been used. The novelty lies in the incorporation of MODIS-derived predictors to foster new insights for estimating long-term solar energy for any region that does not have atmospheric monitoring systems. More importantly, it can rely on remote sensing data for *GSR* prediction. These models can promote solar energy in remote or regional areas where satellites can be employed for long-term evaluation.

2. Theoretical Background

In this section, only the deep learning models are explained in detail; the theoretical explanation of the ANN [32], RF [33], GBM [34], XGBR [35] and DT [36,37] is all elucidated elsewhere since they are well-known methodologies.

2.1. Objective Model: Deep Learning Approach

Deep learning (DL) is a subfield of data-driven models where the algorithm itself learns the internal representation from raw data to perform a regression or a classification process. This is in contrast to classical methods that require carefully-engineered input features based on domain expertise. The DL algorithm can be classified with artificial neural networks because of its multi-layer structure formed by input and output neurons. The multi-layer network as a class of data-driven methods is built by attaching multiple layers to form a unique machine. The DL methodology aims to sequence independent machines, in which the output of one layer is the input of the next layer.

In this paper, we adopt DL as it has recently been used to predict renewable energy sources. For example, the work in [38] used deep belief network (DBN) to predict wind power, whereas [39] applied stacked auto encoders to predict short-term wind speed. Hence, two fundamental forms of DL, based on DNN and DBN, are used to predict *GSR* by employing MODIS-derived predictor variables.

2.1.1. Deep Belief Network

Deep belief network (DBN) is a generative model with a stacked restricted Boltzmann machine (RBM) and a sigmoid belief network. A typical DBN algorithm flowchart is shown in Figure 1a. The deep belief network plays an important role in modelling time series data [23] and has been adopted in energy studies, for example wind speed prediction [40,41].

The proposed DBN *GSR* prediction model is comprised of two RBMs and one MLP (Figure 1b). The RBM is symmetrical bipartite with two layers (visible and hidden layer). The visible units $v = \{v_1, v_2, \cdots, v_m\}$ and the hidden units $h = \{h_1, h_2, \cdots, h_n\}$ are connected by a symmetrical weight matrix W, as well as bias weights (offsets) $a = \{a_i \mid i = 1, 2, \cdots, m\}$ for the visible units and $b = \{b_j \mid j = 1, 2, \cdots, n\}$ for the hidden units. RBM is an energy model, and the energy function of the visible layer and hidden layer is defined as below [42]:

$$E(v,h;\theta) = -\sum_{i=1}^{n_v} a_i v_i - \sum_{j=1}^{n_h} b_j h_j - \sum_{i=1}^{n_v} \sum_{j=1}^{n_h} h_j W_{j,i} v_i$$

where

v_i = number of neurons in the visible layer;

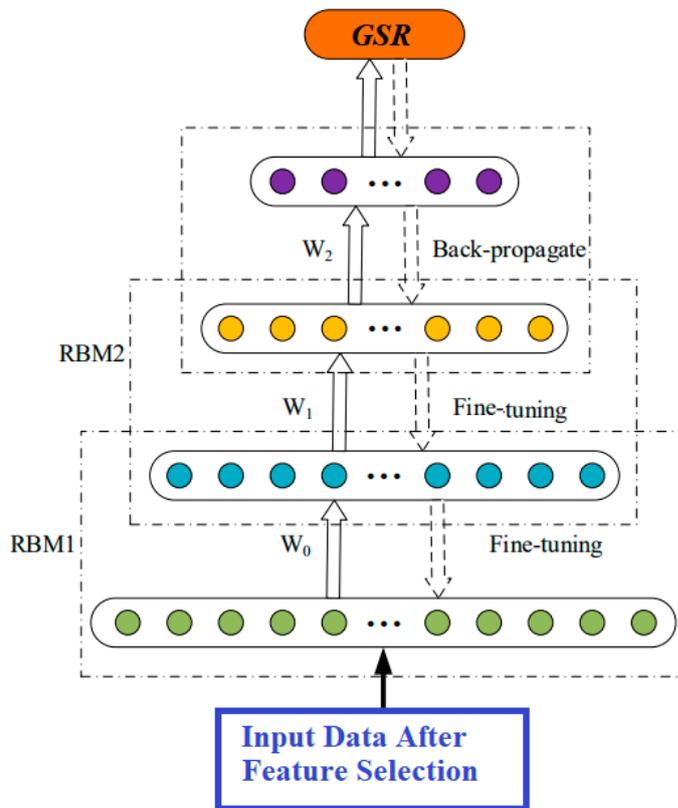a_i = is the bias of node i ;

b_j = is the bias of node j ;   (1)

h_j = is the number of Boolean hidden neurons

    within the hidden layer;

$W_{j,i}$ = is the weight matrices between the

    visible layer and hidden layer;

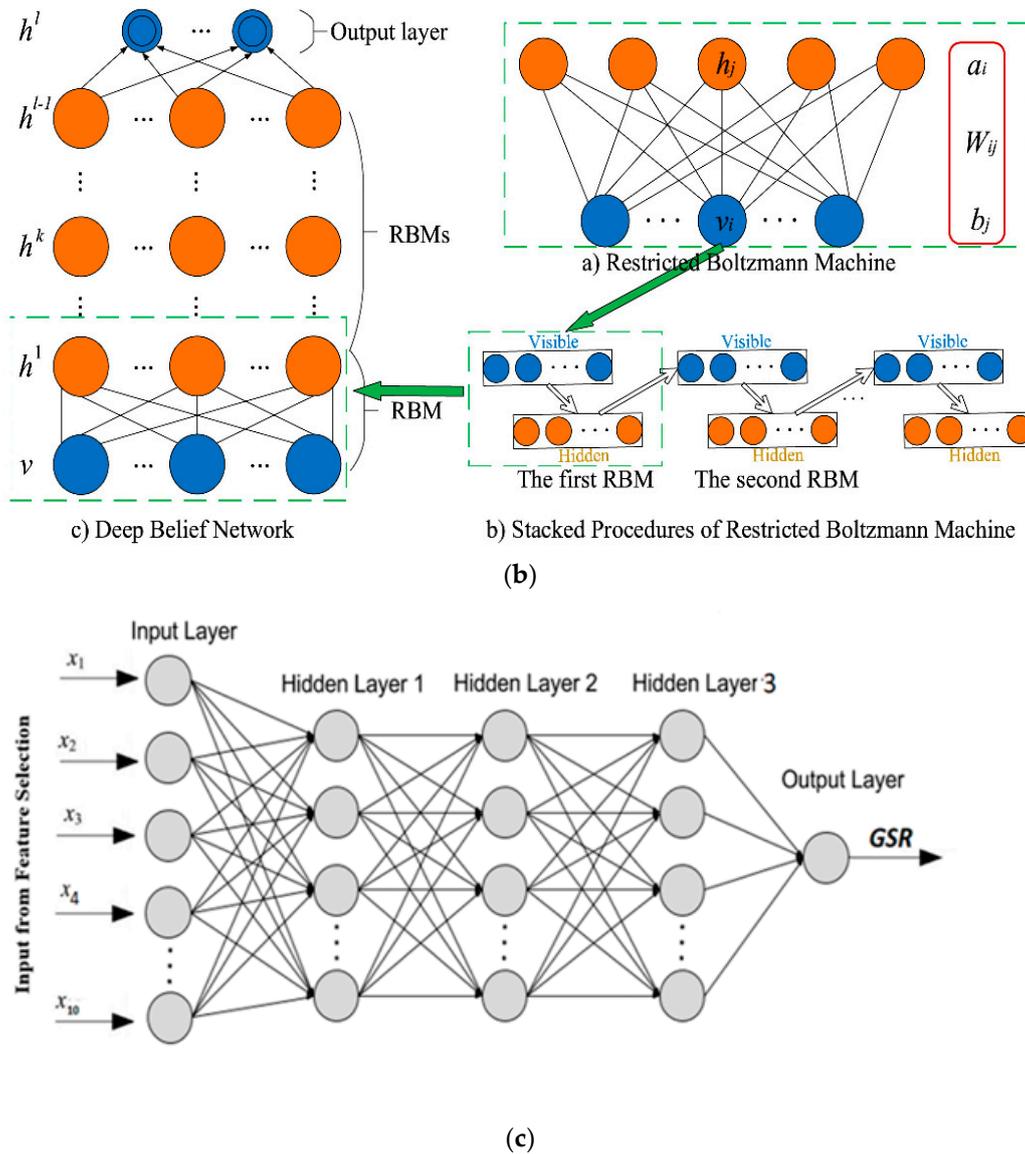θ = is a model parameter, $\theta = \{w,a,b\}$;

(**a**)

(**b**)

(**c**)

Figure 1. (**a**) Deep belief network algorithm flow chart. (**b**) The procedures to implement the deep belief network. (**c**) Topological structure of the deep neural network (DNN). RBM, restricted Boltzmann machine; *GSR*, global solar radiation.

In order to minimize the energy function, i.e., Equation (1), the model parameters $\theta = \{W, a, b\}$ of RBM need to be updated by the contrastive divergence algorithm proposed by Hinton [43], and the update rule can be derived by Equation (2) [44].

$$\Delta \mathbf{W} = \varepsilon(\mathrm{v}\mathrm{h}^T - \mathbf{v}'\mathbf{h}'^T)$$
$$\Delta \mathbf{a} = \varepsilon(\mathrm{v} - \mathrm{v}')$$
$$\Delta \mathbf{b} = \varepsilon(\mathbf{h} - \mathbf{h}')$$
(2)

where ε is the learning rate and V' and h' are the reconstruction of **v** and **h** by Gibbs sampling [45], respectively. Once the first RBM is trained, its hidden layer becomes the visible layer of the next RBM, and the new RBM is trained with the procedure above. Then, a supervised learner MLP is added to the top of the network for time series forecasting. Finally, the parameters of the whole network are fine-tuned by the back-propagation algorithm.

2.1.2. Deep Neural Network

Deep neural networks (DNNs) are complex, yet fully-connected ANNs composed of more than one hidden layer (Figure 1c), where each successive layer uses the outputs from the previous layer. Although different architectures are available, a common DL has a feed-forward network, with a back-propagation algorithm, for the learning and optimization. Although DNNs exhibit superior performance, overfitting is the major issue, which can be decreased by applying a regularization technique such as a weight penalty, early stopping or dropout during training. The input layer of the DNN implemented in this study is selected using the feature selection procedure (Section 3) where one neuron in the output layer is used to generate the predicted *GSR*. The mathematical form of the neural network forward propagation model is described below [46]:

$$a^l_{\ i} = f\left(\sum_{j=1}^{N^{l-1}} W_{ij}^{\ l} a_j^{\ l-1} + b_i^{\ l} \right)$$

where

$a^l_{\ i}$ = output value from the ith neurons in the lth layer neural network.

$a_j^{\ l-1}$ = output value from the ith neurons in the $(l-1)^{th}$ layer neural network

$W_{ij}^{\ l}$ = the weight from the jth neurons in the $(l-1)^{th}$ layer to the ith neurons in the lth layer. (3)

$b_i^{\ l}$ = bias term from the ith neurons in the lth layer neural network.

N^{l-1} = number of neurons in the $(l-1)^{th}$ layer.

$f(\bullet)$ = is the activation function of the neurons.?

Among the many types of neural network activation functions, a popular one includes the sigmoid function, ReLU (rectified linear unit), softplus and hyperbolic tangent (tanh) [47] functions, described as follows:

$$f(x) = \frac{1}{1+e^{-x}} = sigmoid\ function\ (\sigma)$$

$$f(x) = \frac{e^{2x}-1}{e^{2x}+1} = tanh function$$

$$f(x) = ln(e^x+1) - ln2 = softplus\ function$$ (4)

$$f(x) = max(0,x) = ReLU\ function$$

where x is the input to a neuron

The mean squared error of training sets can be described by [32]:

$$mse = \frac{1}{m} \sum_{t=1}^{m} (o_t - y_t)^2 = T(w)$$

where

mse = mean squared error between the predicted and true values of the data sets.

$o_t = t^{th}$ sample's predicted value of *GSR*. (5)

$y_t = t^{th}$ sample's actual value of *GSR*

w = vector that contains the weights and bias terms between the neurons in each layer.

m = number of the datasets.

This paper uses the adaptive moment estimation (Adam) [48], root mean squared prop (RMSProp) [49], adaptive gradient (AdaGrad) [50], Nesterov-accelerated adaptive moment estimation (Nadam) [51], the variant of Adam based on the infinity norm (Adamax) [48] and the

adaptive delta (AdaDelta) [52] algorithm to avoid the model from falling into a local optimal solution. These algorithms applied updated weights to achieve high efficiency and fast convergence. More details regarding the learning algorithm are found in other works [53].

3. Data, Importance and Context of the Study

This study employs monthly averaged daily *GSR* records to develop a prediction model using DBN and DNN for four solar cities of Australia: Blacktown [33.77°S, 150.90°E], Adelaide [34.92°S, 138.59°E], Townsville [19.25°S, 146.81°E] and Central Victoria [36.74°S, 144.28°E]. Although the potential for use of solar energy in these regions remains high, deep learning-based models for *GSR* are not easily available. Furthermore, in most states in Australia, the electricity is provided through power plants located in the central and southern areas, and because of this, there are huge transmission and distribution costs and losses [54]. Hence, there is a potential to harness locally available solar energy, particularly in remote sites where solar forecast models are actively being tested [9–13], although the studies are using simplistic models rather than deep learning approaches.

Other than focusing on solar city sites with abundant solar radiation, this study purposely adopts MODIS satellite variables to model *GSR* since historical data related to the target variable (*GSR*) play a key role in helping evaluate solar energy availability. Remote sensing data have already been identified as a practical predictor for solar problems [55], so in this view, the coupling of a deep learning model with satellite-derived products is a major improvement over the use of station-based data mainly because the acquisition of satellite imagery can be feasible for inaccessible sites with no measurement infrastructure as long as a footprint is identified. For long-term forecast horizons (*e.g.*, monthly), satellite data remain abundant for a diverse range of spatial and temporal resolutions and, recently, have been adopted in global solar radiation prediction problems [9,13]. Although recent studies have considered solar radiation models trained with MODIS datasets, these were limited to cloud-free predictor variables and land surface temperature. Considering this, significant MODIS data have not been used in previous studies [56], although a recent study [12] has estimated solar radiation using MODIS-derived predictors without a deep learning model.

3.1. MODIS Satellite-Derived Predictor Data

To design a deep learning model for *GSR* prediction over long time horizons, monthly predictor data have been extracted from 1 March 2000–2018 from NASA's Goddard Online Interactive Visualization and Analysis Infrastructure (GIOVANNI) repository. Table 1 lists the predictors. The objective variable (*i.e.*, integer values of land surface daily global solar radiation) were downloaded from a ground-based source, the Scientific Information for Land Owners (SILO) database. The Long Paddock SILO database is operated by Queensland Government Department of Environment and Science in the Department of Science, Information Technology, Innovation and the Arts (DSITIA). This data cover each of the four solar cities [57].

Table 1. Description of Moderate Resolution Imaging Spectroradiometer (MODIS) satellite-derived predictors, with the relevant notation adopted in this study to predict monthly averaged daily solar radiation (*GSR*) in Australia's solar cities (data source: Goddard Online Interactive Visualization and Analysis Infrastructure (GIOVANNI) NASA Repository).

Data Source	MODIS-Derived Variable	Notation	Units
	Aerosol Optical Depth (550) Dark Target Deep Blue Combined	*aoddtdbc*	none
	Aerosol Optical Depth Land Ocean	*aodlc*	none
	Aerosol Scattering Angle	*asa*	none
	Atmospheric Water Vapour Medium	*awvm*	cm
	Atmospheric Water Vapour High	*awvh*	cm
	Atmospheric Water Vapour Low	*awvl*	cm
	Cloud Effective Radius Ice	*cefri*	μm
	Cloud Effective Radius Liquid	*cerl*	μm
	Cloud Fraction	*cf*	none
	Cloud Fraction Day	*cfd*	none
	Cloud Fraction Night	*cfn*	none
	Cloud Optical Thickness Combined	*cotc*	none
GIOVANNI (MODIS Level 3 Atmosphere Products: MOD08_M3)	Cloud Optical Thickness Ice	*coti*	none
	Cloud Optical Thickness Liquid	*cotl*	none
	Cirrus Reflectance	*cr*	none
	Cloud Top Pressure Night	*ctp*	hPa
	Cloud Top Pressure	*ctpd*	hPa
	Cloud Top Pressure Day	*ctpd*	hPa
	Cloud Top Temperature	*ctt*	K
	Cloud Top Temperature Day	*cttd*	K
	Cloud Top Temperature Night	*cttn*	K
	Cloud Water Path Ice	*cwpi*	gm^2
	Cloud Water Path Liquid	*cwpl*	gm^2
	Deep Blue Angstrom Exponent Land	*dbael*	none
	Deep Blue Aerosol Optical Depth 550 Land	*dbaodl*	none
	Water Vapour Near Infrared Clear	*wvnic*	cm
	Water Vapour Near Infrared Cloud	*wvnicl*	cm

The GIOVANNI data offer a fast and flexible method to explore links between physical, chemical and biological parameters useful for inter-comparing multiple satellite sensors and algorithms [58]. Since only a relatively short investment of time and effort is required to become familiar with the GIOVANNI system, a main advantage is its ease-of-use so that researchers who are unfamiliar with remote sensing can use the system to determine their data needs applicable to their topic area [59]. Missions, instruments, or projects providing data products available in GIOVANNI are useful for *GSR* modelling as they include the Atmospheric Infrared Sounder (AIRS), Tropical Rainfall Measuring Mission (TRMM), Ozone Measuring Instrument (OMI), Moderate Resolution Imaging Spectroradiometer (MODIS), Modern Era Retrospective-analysis for Research and Applications (MERRA) project and North American Land Data.

In this paper, data from a MODIS instrument on-board Terra (EOS AM) and Aqua (EOS PM) satellites have been utilized. These satellite (MODIS) meteorological data are widely and freely

available for public access [60] and, therefore, useful for solar energy exploration and modelling in a diverse range of sites.

3.2. Data Preparation, Feature Selection and Sensitivity Analysis

Before the *GSR* model was developed, all inputs were normalized in the range of (0, 1) [9]. Normalization was done to have the same range of values for each of the inputs to the models. This normalization procedure guarantees stable convergence of weight and biases [61,62].

$$X_n = \frac{X_{actual} - X_{min}}{X_{max} - X_{min}} \tag{6}$$

where X, X_{min} and X_{max} represent input data and minimum and maximum values, respectively.

After this, the data were segregated into training and testing sets. Since there is no rule for data segregation, we used an earlier researcher's approach [63,64] to divide into 80% (training) and 20% (testing) sub-sets, but 10% of the training data were separated again for the purpose of model validation, mainly to eliminate issues related to a model bias through a cross-validation process.

In this paper, a total of five filter- and 10 wrapper-based feature selection (FS) algorithms were employed to extract the most important MODIS-derived predictors related to the target (*i.e., GSR*). Table 2a,b outlines the FS algorithms. By removing irrelevant, noisy or redundant features from the original space, FS can alleviate the problem of overfitting, improve the performance [65] and save time and space costs that are normally an issue of consideration in a deep learning algorithm [66]. Importantly, through an FS strategy, we can also get deeper insights into the MODIS and *GSR* data by analysing the importance of all and the most relevant features that can affect the future sustainability of solar energy.

For this study, FS divided into two categories, filters and wrappers, has been used. The filter method was used as a pre-processing step using criteria that did not involve any learning, and by doing that, it did not consider the effect of a selected feature subset on the performance of the algorithm [67,68]. Wrapper methods, on the other hand, were used to evaluate a subset of features according to the accuracy of a predictor [65], where search strategies were used to yield nested subsets of variables, and the variable selection was based on the performance of a learned algorithm [69]. In accordance with Table 2a,b, this study had multiple FS algorithms to select the most optimal predictors of long-term *GSR* carefully.

Table 2(a). Description of the 15 feature selection algorithms applied to obtain the best predictors of GSR from a global pool of MODIS-derived variables used to predict long-term *GSR* in Australia's solar cities.

Name of Feature Selection Algorithm	Notation	Feature Extraction Method
Particle Swarm Optimization	PSO	Wrapper
Genetic Algorithm	GA	Wrapper
Simulated Annealing	SA	Wrapper
Stepwise Regression	Step	Filter
Nearest Component Analysis Regression	FSRNCA	Wrapper
Relief Algorithm	Relief	Filter
Ant Colony Optimization	ACO	Wrapper
Nondominated Sorting Genetic Algorithm	NSGA	Wrapper
Random Forest Regressor	RF	Wrapper
Univariate Feature	UNV	Filter
Exhaustive Search	EXH	Wrapper
Mutual Information Regression	MIR	Filter
Sequential Backward Selection	SBR	Wrapper
Sequential Forward Selection	SFR	Wrapper
Recursive Feature Elimination	RFER	Wrapper

1

Table 2(b). List of MODIS-derived predictors screened at each solar city in Australia after applying the feature selection algorithm (Table 2a). All notations as per Tables 1 and 2a.

Solar City Location	SFR	SBR	RFER	MIR	EXH	UNV	RF	ACO	FSRNCA	GA	PSO	NSGA	Step	Relief	SA
Adelaide	aod	asa	awvl	aod	aod	asa	asa	asa	asa	asa	asa	asa	asa	aod	asa
	asa	cfd	awvh	asa	asa	awvh	awvl	awvl	awvh	awvh	awvl	awvh	cfd	asa	awvh
	cfd	cfn	cfd	awvl	aod	awvl	awvm	cerl	awvh	cfd	awvm	awvl	cotc	aod	cfd
	cfn	cotl	cerl	awvm	awvl	awvm	cfd	cfd	cfd	cfm	ceri	awvm		awvh	cfm
	cotc	ctpd	aod	cerl	awvm	cerl	dbael	cfm	cfm	cfn	ceri	cfd		awvl	cfn
	ctpd	ctpm	asa	cfd	awvh	cfd		ctpd	cfn	cotc	cfd	cfm		awvm	cotc
	ctpm	ctpn		cfm	cerl	cfm		ctpm	cotc	coti	cfn	cotc		cerl	coti
	cttd	cttm		dbael	cfd	dbael		cttm	coti	cttm	cotc	ctpd		cfd	cttm
		dbael		wvnic	cfm	wvnic		cttn	cttd	cttn	cotl	ctpn		cfm	cttn
		wvnic			dbael	wvnicl		wvnicl	cttn	dbaod	cttd	cttm		dbael	wvnic
					wvnic				wvnicl	wvnic	cttm	wvnicl			
Townsville	cttm	aod	cttm	asa	aod	aod	asa	aod	asa	asa	asa	aod	asa	asa	cfd
	wvnicl	asa	wvnicl	cerl	asa	asa	cerl	cttn	awvl	awvh	awvm	asa	awvh	awvl	ctpm
	cfn	awvh	cfm	cfd	awvh	aod	cfd	awvl	awvh	awvl	awvl	awvl		awvm	asa
	dbael	ceri	dbael	cfm	awvh	awvh	dbael	asa	ceri	awvm	cfd	awvm		awvh	awvl
	dbaod	cfd	dbaod	cfn	ceri	cerl		cttm	cerl	ceri	coti	cfd		cfd	cotc
	cfd	cotc	cfd	cotl	cfd	cfd		cotl	cfm	cerl	cotl	cfm		cfm	awvh
	cotc	coti	cotc	ctpd	cotl	cotc		cfm	cotc	cfd	ctpm	cotc		cfn	dbaod
	aod	cotl	aod	ctpn	ctpd	cotl		dbael	ctpm	cfm	cttm	cotl		cttd	ceri
	asa	ctpd	asa	cttd	dbael	cttd		cerl	dbaod	cfn	cttn	ctpd		cttm	ctpn
		dbael		dbael	wvnicl				dbael	dbael	cotc	dbael		wvnic	cttn
										cotc					
										cotl					
										ctpm					
Blacktown	asa	asa	dbael	asa	aod	asa	asa	aod	asa	asa	aod	asa	asa	aod	aod
	awvh	awvh	ctpd	awvh	asa	awvh	awvl	asa	awvh	awvl	awvl	awvl	cfd	asa	asa
	awvl	awvl	cttn	awvl	awvh	awvl	awvm	cfd	ceri	awvm	cfd	awvm		aod	awvh
	cfd	cfd	cfn	awvm	cr	awvm	dbael	cfm	cerl	cfd	cfm	ceri		awvl	ceri
	coti	cotc	coti	cerl	ceri	cerl		cfn	cfd	cotc	cotc	cfd		awvm	cfd
	cttd	cotl	awvh	cfd	cfd	cotl		cotc	cfm	coti	ctpm	coti		cerl	cotc
	cttm	ctpd	cotc	cfm	cotc	dbael		ctpd	cfn	cotl	cttd	cotl		cfd	coti
	cttn	ctpm	cfd	ctpd	coti	wvnic		ctpm	cotc	ctpd	cttn	ctpn		coti	cotl

	dbaod	cttm	aod	dbael	ctpd	aod		cttn	coti	ctpm	wvnicl	wvnic		dbael	cttm
	wvnicl	cttn	asa	wvnic	cttd				ctpd	cttn		wvnicl		wvnicl	wvnicl
										wvnicl					
Central Victoria	asa	aod	cttm	asa	asa	asa	asa	asa	asa	asa	asa	aod	asa	aod	asa
	awvl	asa	ctpm	awvh	awvh	awvh	awvl	dtdb	dtdb	awvh	awvh	asa	cfm	asa	awvl
	cfd	cr	cttd	awvm	cr	awvl	awvm	awvl	ceri	awvl	awvl	awvl		dtdb	awvm
	cfn	cfd	cotc	cfd	ceri	awvm	cfd	awvm	cfd	awvm	awvm	awvm		awvh	cr
	coti	cfn	wvnic	cfm	cerl	cfd	dbael	cr	cfm	ceri	cfm	cr		awvl	cerl
	ctpm	coti	awvm	ctpd	cfd	cfm		cfd	coti	cerl	cfn	ceri		awvm	cfd
	cttd	ctpd	dtdb	dbaod	coti	ctpd		cfm	ctpd	cfd	cfn	cfd		cfd	cotc
	cttm	cttm	ctpd	dbael	ctpd	ctpm		cotc	cttn	cfm	ctpm	cfm		cfm	ctpd
	dbael	cttn	awvh	wvnic	cttm	dbael		dbael	dbaod	coti	cttn	ctpd		dbael	dbael
			cfd		dbaod	wvnic		wvnic	dbael	cotl	cttm	dbael		wvnic	wvnic
			cr							ctpd	dbael	wvnicl			
			ceri							dbael					

Other than incorporating the FS strategy, we also performed a sensitivity test to examine the statistical relationships between *GSR* and its selected variables. To estimate *GSR* in a region with limited predictors, a solar engineer may be interested in checking the importance of a given set of predictors that effectively contribute to a predictive model. This information is useful for decision making in solar power plant design, especially in selecting the most appropriate predictors for *GSR* and enhancing the understanding of the correct measurements to obtain when those data are used. In this study, we employed a global sensitivity analysis method using the Gaussian Emulation Machine (GEM-SA) software [70]. For detailed information on this technique, readers can consult [71]. To deduce which of the MODIS-derived inputs produced a substantial effect on the target variable (*GSR*), two GEM-SA parameters were used: the main effect (*ME*) and the total effect (*TE*). The *ME* enumerates the influence of just one parameter varying in its own in relation to *GSR*, while the *TE* comprises the *ME* plus any variance due to possible interactions between that parameter and all of the other inputs varying at the same time [72].

Figure 2 is a "Lowry plot" and shows the relative contribution to the total variance in *GSR*, from each selected MODIS input. Notably, the vertical bars show *ME* and *TE* for each input ranked in order of their main importance, whilst the lower and upper bounds show the cumulative sum of the main and total effect, respectively. This analysis shows that almost 72%, 50% and 27% of the variance in *GSR* was due to the *asa* variable (*i.e.*, aerosol scattering angle) for the Adelaide, Blacktown and Townsville study sites, respectively. In contrast, for the case of Central Victoria, almost 40% of the total variance in *GSR* was due to *awvm* (*i.e.*, medium atmospheric water vapour) compared to about 23% due to *asa*. For Adelaide, however, the second highest contribution was derived from day-time cloud fraction (*cfd* ≈ 27%). It can therefore be concluded that for Adelaide, these input variables are important and are likely to affect the performance of the deep learning models if they are neglected. Similarly, for Central Victoria, the low atmospheric water vapour (*awvl*), aerosol scattering angle (*asa*), and day-time cloud fraction (*cfd*) were found to be the second, third and fourth highest contributors analysed by the GEM-SA method, whereas the other MODIS input variables appeared to have a negligible effect (<5%).

In contrast to the above results, for the case of Blacktown, all of the other MODIS-derived inputs had a negligible effect on *GSR* with less than 10% of the total variance. It can therefore be concluded that to include 90% of the variance, the first six MODIS parameters (*asa, awvl, awvm, cfd and cttn*) are required in modelling *GSR* for Blacktown. Similarly, the first 10 MODIS parameters (*asa, awvm cfm, awvl, cfd, ctpm, awvh, cotc and dbael*) are required for Townsville to include 90% of the variance. The effect of MODIS inputs on the target variable can be easily identified with the GEM-SA method. As revealed in this analysis, it should be noted that the most important MODIS inputs are not the same for all four locations; hence, a sensitivity analysis of FS-based inputs is necessary to identify more clearly the role of these predictors in modelling the objective variable.
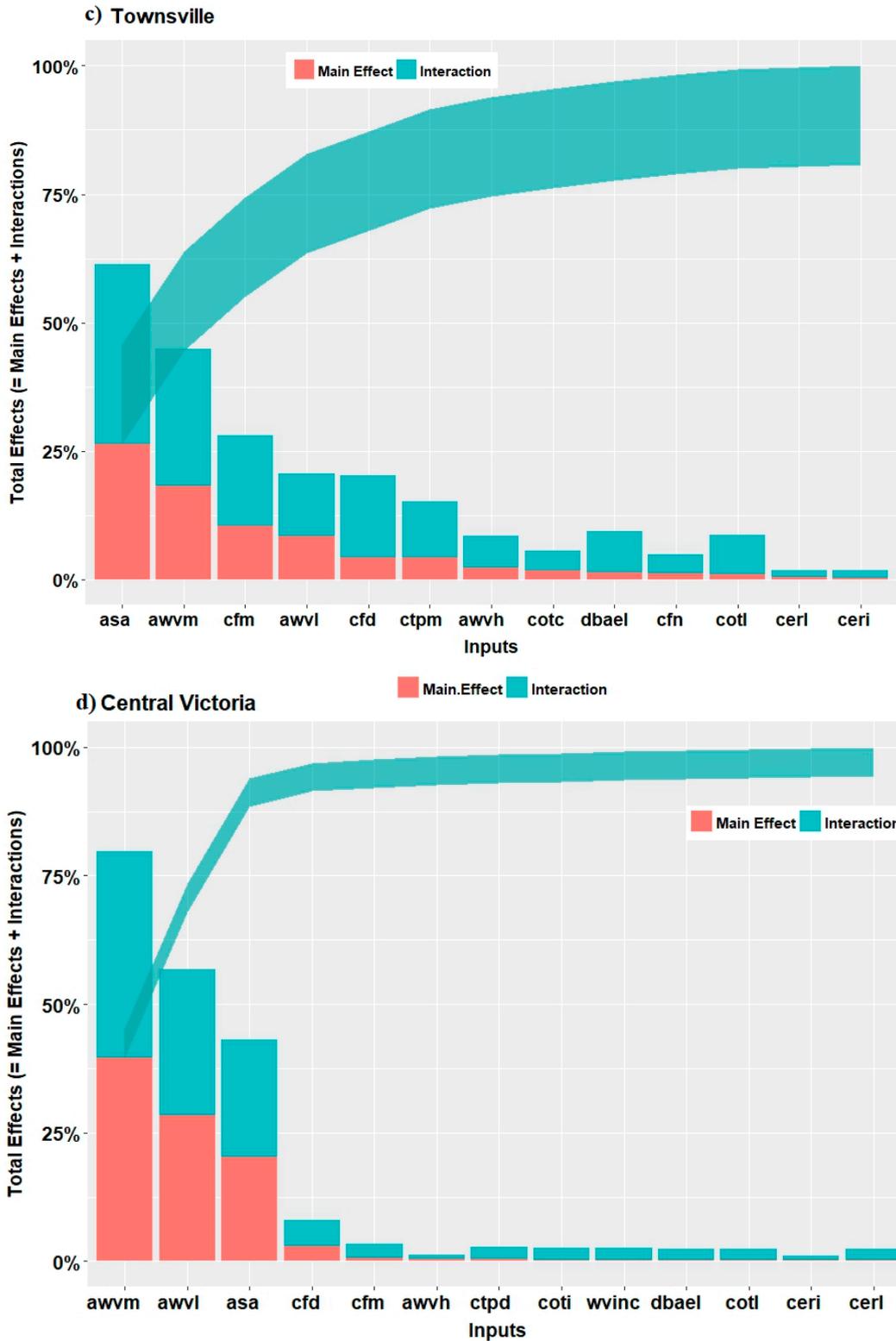
a) Blacktown

b) Adelaide
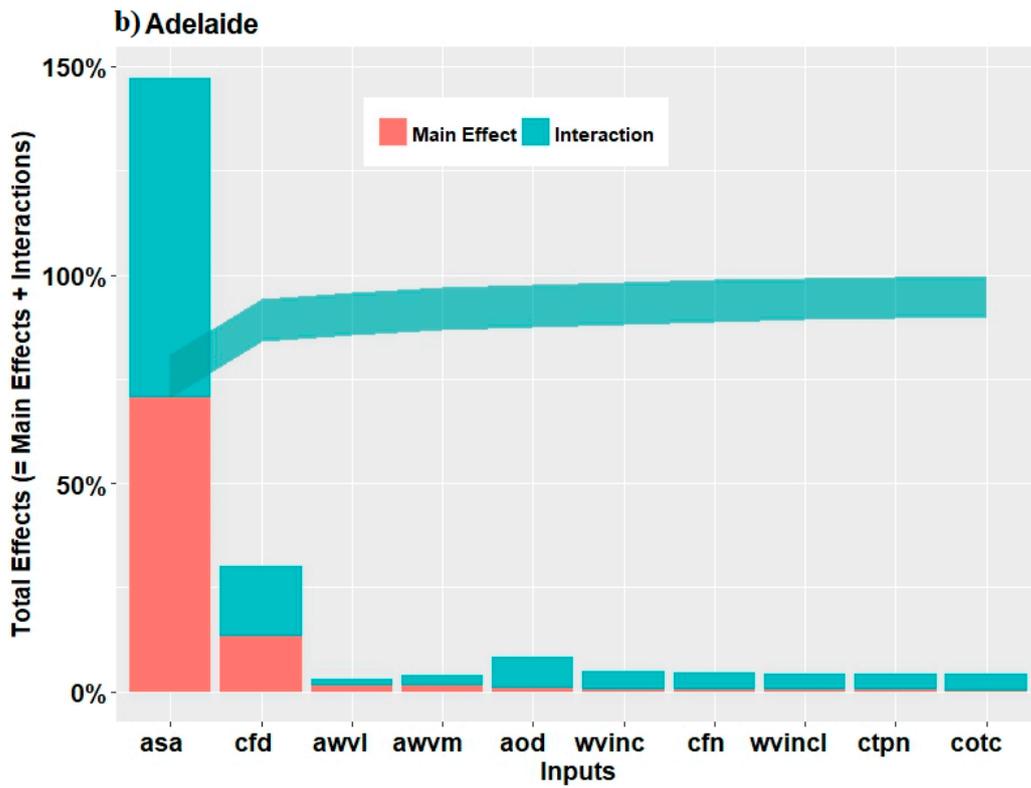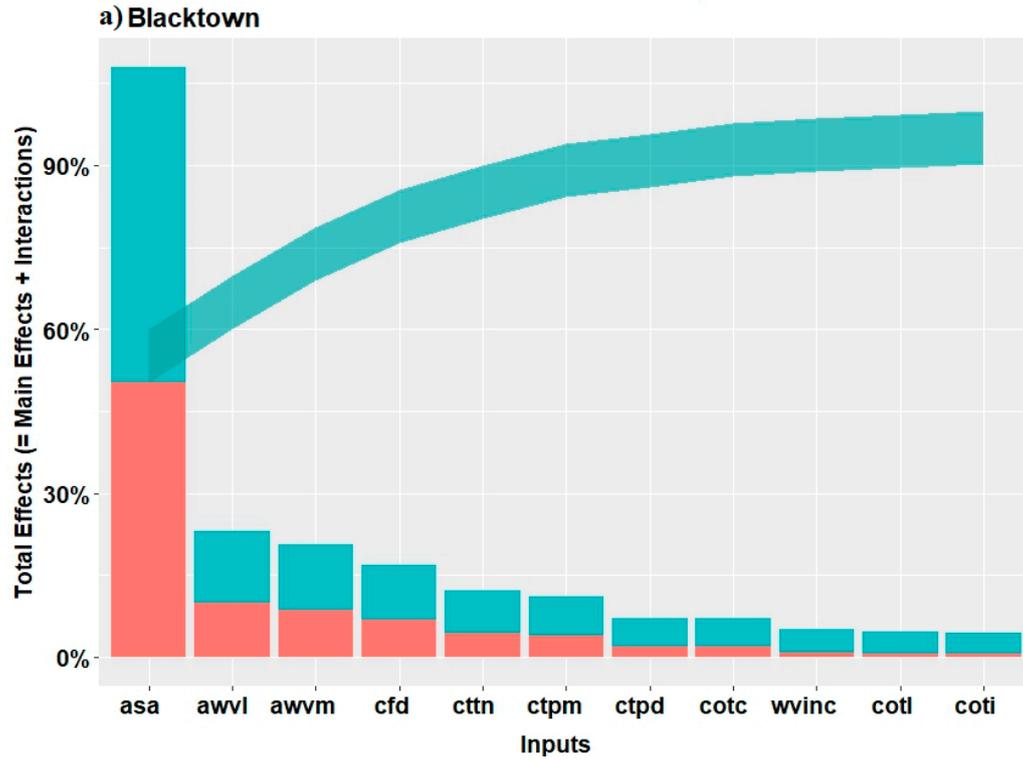
Figure 2. Lowry plot with the main (primary) and the total cumulative effects of MODIS-derived predictors employed to predict monthly averaged daily global solar radiation *GSR*.

3.3. Deep Learning Predictive Model Design

In this study, deep learning was implemented in Python with the Keras Deep Learning library together with Theano [73] used for modelling *GSR* in a computer with an Intel core *i7* processor @ 3.3 GHz and with 16 GB RAM memory.

3.3.1. Deep Belief Networks

After a feature selection process and sensitivity analysis of MODIS-derived predictors, a DBN model architecture was designed. This study followed the notion that there is no theoretical basis to set a correct number of layers in a deep learning model. Indeed, insufficient hidden layers means that there could be no proper feature space, resulting in an under-fitted model, but too many layers can lead to the issues of over-fitting, as well as an "ill-posed" problem with higher computational costs [74]. Considering these, a trial and error method was adopted to determine the optimal structure of a DBN model, selected carefully from a total of 12 different neuronal architectures.

For the DBN model, this study used back-propagation for all trained models, but the activation functions were switched between rectified linear unit (ReLU) and the sigmoid equation with a regularization parameter used for fine tuning. The finer details of DBM models are as follows.

(1) Back-propagation was used to adjust weights, using the derivative chain principle on model errors that were propagated from the last to the first layer. The two parameters implemented were the batch size [2,5] and epochs (100, 200), where training samples were divided into groups of the same size. Notably, the batch size refers to the samples in each group fed to the network before weight updates are performed, whereas epochs are related to the iterations of fine-tuning. Generally, the network can undergo fine-tuning with a smaller batch size or a larger number of epochs [75] including a large iteration set of 1000 in this study.

(2) To avoid overfitting, a least absolute shrinkage and selection operator (*i.e.*, *L2 or lasso* regularization) was used to update the cost function by adding a regularization term [76], such that the weights were reduced, to assume a neural network with a smaller weight matrix, leading to a cost-efficient DBN model. This is likely to reduce the overfitting [77], so in this study, we used the *L2* regularization as 0.01.

(3) The learning rate for stacked restricted Boltzmann machine (RBM) and back-propagation were fixed to 0.01 and 0.001 for the DBN model design following earlier studies [78].

In accordance with Table 2, the input nodes were selected by the feature selection algorithm, and hidden layers were deduced by trial and error with analysis of the influence on training performance. As a result, one hidden layer was used at first, and increased up to two layers with variable layers and neurons to optimize the predictive model. This resulted in 12 distinct DBN architectures where the DBN_{10} model was the optimal model.

Table 3 lists the effect of feature selection in designing the optimal DBN model, where the relative root mean square (*RRMSE %*) generated for a selected study site, Adelaide in the model training phase, is illustrated. Evidently, MODIS-based predictors acquired through the particle swarm optimization (PSO) algorithm yielded the lowest *RRMSE* (≈ 2.98%) in the training DBN_{10} model, as identified in Table 3.

Table 3. The influence of feature selection algorithms on *GSR* prediction problems in terms of the relative root mean square (*RRMSE* %) generated by the deep belief network (DBN) model for a selected solar city, Adelaide (Australia) in the training phase illustrated as an example. The most optimal feature selection algorithm (*i.e.*, PSO) and the relevant DBN model architectures (*i.e.*, DBN_{10}) are highlighted in blue and is **boldface**.

Feature Selection Algorithm	DBN_1	DBN_2	DBN_3	DBN_4	DBN_5	DBN_6	DBN_7	DBN_8	DBN_9	$\mathbf{DBN_{10}}$	DBN_{11}	DBN_{12}
ACO	3.8032	3.605	6.0055	4.1061	3.8771	5.9773	3.4115	3.6646	3.7953	4.9275	4.0954	4.9497
EXH	4.6274	4.4285	4.1359	5.7535	6.1618	5.1044	5.4939	6.0681	5.9135	4.4656	6.469	5.4435
FSRNCA	3.7848	4.3611	3.3554	3.2943	4.3768	3.8191	3.7903	5.3001	3.4902	4.0396	4.282	4.9547
GA	5.1471	3.2947	3.5085	3.8677	4.0935	5.5466	3.8229	6.5333	3.8884	3.4569	3.6279	4.0753
MIR	4.7199	4.3303	4.5359	4.3908	5.1334	4.7249	4.969	9.4603	4.8635	4.3226	6.7647	4.5045
NSGA	3.2782	3.3948	4.4819	5.2861	4.1812	3.3408	4.9873	7.3162	4.4266	3.431	3.4871	3.9187
PSO	2.9939	4.9381	8.214	3.0133	4.8548	4.3518	4.6447	3.5864	3.6208	**2.9888**	4.512	3.2131
Relief	4.7294	4.7415	4.5825	6.0117	5.516	6.0632	4.8268	5.2409	4.3927	4.4404	4.7927	7.5023
RFER	4.2767	3.9111	3.9301	3.8939	4.5415	4.8686	4.4713	5.6978	4.3635	3.9684	4.4727	4.3049
RF	4.4939	5.4009	4.2472	4.293	6.3196	4.6522	4.3344	4.4513	4.678	4.4939	5.1191	4.5643
SA	3.29	3.3942	3.3791	4.0482	4.5882	3.4462	3.7251	3.4854	3.5494	4.0958	4.0854	3.3233
SBR	4.4648	4.339	6.6621	5.3684	4.3839	4.0238	4.1885	4.6828	3.9967	4.4697	5.0949	6.799
SFR	3.4398	3.3889	4.8542	3.7915	4.8091	3.7299	4.4864	4.4268	3.7121	3.1502	4.1725	3.4217
Step	3.4244	3.3991	3.6212	3.3992	4.4455	3.8746	4.0521	3.3514	4.5755	3.457	3.8704	3.6596
UNV	4.6064	5.0044	4.7281	4.293	5.7903	4.6687	5.2329	5.6587	4.7297	4.6151	4.7372	5.0933

Similarly (not shown here), the MODIS-derived predictors analysed by the genetic algorithm with DBN$_{11}$ (*RRMSE* ≈ 3.25%), analysed with step feature selection for DBN$_2$ (*RRMSE* ≈ 3.79%) and the relief algorithm with DBN$_2$ (*RRMSE* ≈ 3.71%), yielded the lowest *RRMSE* compared to the other DBN models for Blacktown, Townsville and Central Victoria, respectively. In addition, the increase in neurons in the hidden layers above 50 was seen to increase training errors for Adelaide, with *RRMSE* being elevated by 65%, 174%, 81%, 62%, 45%, 55%, 19% and 21% for DBN$_2$, DBN$_3$, DBN$_4$, DBN$_5$, DBN$_6$, DBN$_7$, DBN$_8$ and DBN$_9$, respectively (not shown here). In this study, a total of 180 DBN architectures were developed to generate the optimal *GSR* predictive model.

3.3.2. Deep Neural Network

To design a competing deep learning approach for *GSR* prediction, the next objective model, DNN with 3 hidden layers, 1 input layer where MODIS-derived predictor variables were incorporated from the FS process and 1 output layer corresponding to the target (*i.e.*, *GSR*), was designed. As with the case of DBN (Section 3.3.1, there appeared to be no preferred method to optimize a DNN model, so a trial and error approach was implemented, where the number of neurons in the hidden layer, activation function, batch size, and number of neurons were tested randomly to satisfy the most accurate *GSR* model.

Specifically, the modelling experiments were executed 10 times for the same DNN configuration to attain the best result, with the various steps as follows.

- In each trial and error, DNN was trained using popular algorithms: AdaGrad, RMSProp, AdaDelta, Adam, Adamax, Nadam and *SGD*. It is noteworthy that the Adam algorithm is normally quite popular [79], given it is an enhanced combination of *RMSProp* and the *moments* techniques [80]. In this study, we utilized all seven algorithms to determine the optimal DNN architecture.
- To avoid overfitting, three measures were employed. First, we added *L2* regularization to penalize the weights in the deep neural network. Second, the dropout technique was used to omit the subset of hidden units at each iteration of a training procedure [81]. Third, early stopping was applied by monitoring the validation performance with the last 10% of the training dataset, in accord with earlier studies [82].

In total, six distinct DNN model architectures with different hyperparameters were developed. Table 4 shows the architecture for one study site (Central Victoria), where the best model designed with the *Adam* and *SGD* algorithms was shown to generate the lowest *RRMSE*.

Table 4. The architecture of 6 different DNN designed with the back-propagation algorithm for GSR prediction.

				Architecture of DNN						
Model	**Hidden Layer 1 (H1)**	**H1 Activation function**	**Dropout percentage**	**Activation function**	**Hidden Layer 2 (H2)**	**H2 Activation function**	**Hidden Layer 3 (H3)**	**H3 Activation function**	**Batch Size**	**Epochs**
DNN_1	500	Sigmoid	0.2	ReLU	200	Sigmoid	500	Sigmoid	1	1000
DNN_2	500	Sigmoid	0.2	ReLU	200	Sigmoid	500	Sigmoid	3	1000
DNN_3	50	Sigmoid	0.2	ReLU	20	Sigmoid	5	Sigmoid	3	1000
DNN_4	500	Sigmoid	0.2	ReLU	50	Sigmoid	20	Sigmoid	3	200
DNN_5	100	ReLU	0.2	ReLU	50	ReLU	20	tanh	1	200
DNN_6	100	ReLU	0.2	ReLU	50	Sigmoid	20	tanh	5	500

Architecture of Back-propagation **(BP) Algorithm for DNN**

BP Optimizers for the DNN Model	**Learning rate**	**Epsilon, ε**	**Decay, δ**	**Rho, ρ**	**Beta, $β_1$**	**Beta, $β_2$**
Gradient-based optimization, AdaGrad	0.01	None	0			
Geoff Hinton's adaptive learning rate method, RMSProp	0.001	None	0	0.9	0.9	0.999
Extended AdaGrad algorithm, AdaDelta	1	None	0	0.95	0.9	0.999
Adaptive Moment Estimation, Adam	0.001	None	0			
Kingma and Ba, (2015) Adamax	0.002	None	0			
Nesterov-accelerated adaptive moment estimation, Nadam	0.002	None	0.004			
Stochastic gradient descent, SGD	0.01	None	0			

where

$\boldsymbol{\delta}$ = Learning rate decay over each update,

$\boldsymbol{\rho}$ = Decay factor

$\boldsymbol{\varepsilon}$ = Factor for updating the variables to eliminate dividing by zero

$\boldsymbol{\beta_1}$ = The exponential decay rate for the first moment estimates

$\boldsymbol{\beta_2}$ = The exponential decay rate for the second moment estimates

Note: ReLU and tanh stand for rectified linear units and hyperbolic tangent activation functions, respectively.

3.3.3. Comparison Models

To benchmark the objective deep learning model (*i.e.*, DBN and DNN), this study used the *Scikit* package [83] to design a Python-based predictive model for *GSR* with XGBoost [84], gradient boosting regression [85], decision tree [86] and the random forest regressor [87]. For the tuning of the regression model's hyperparameters, the study used a grid search [88] package where several parameters like the maximum depth of the tree (*max_depth*), the number of samples required to split an internal node (*min_samples_split*), the number of features to consider when looking for the best split (*max_features*) and the others were tuned.

Table 5a,b shows a full list of parameters tuned by the grid search method with 10-fold cross-validation, where the optimal parameter for each of the study sites, yielding the lowest *RRMSE*, is shown.

For the ANN model, MATLAB 2017b software was utilized [89]. In this study, ANN with various hidden neurons in its hidden layer (1–50) was used with the Levenberg–Marquardt back-propagation algorithm (*trainlm*) [90] including the hyperbolic tangent and logarithmic sigmoid as activation functions, tested in hidden and output layers, respectively. The ANN model with the low *RRMSE* and high correlation coefficient (*r*) was selected.

Table 5(a). Architecture of the decision tree and ensemble-based models developed for *GSR* prediction.

Model	Model Hyperparameters	Acronym	Search Space in Grid Search for Hyperparameter Optimization
Decision Tree	Maximum depth of the tree	*max_depth*	[1,2,5,6,7,8,9,10,11,12,13,14,15,16,17,18,19,20]
	Minimum number of samples to split an internal node	*min_samples_split*	[2, 3, 5, 8,10,12,14,16,18,20,22,24,26]
	Number of features for the best split	*max_features*	['auto', 'sqrt', 'log2']
Random Forest Regressor	Number of trees in the forest	*n_estimators*	[10,20,30,40,50,60,70,80,90,100,120,140,160,180,200,250,300,350,400,450,500]
	Maximum depth of the tree	*max_depth*	[1,2,3,4,5,10,15,20,25,30,35,40,45,50,55,60,80,90,100]
	Minimum number of samples for an internal node	*min_samples_split*	[2, 3, 5, 8]
	Number of features for the best split	*max_features*	['auto', 'sqrt', 'log2']
Gradient Boosting Regressor	Number of boosting stages	*n_estimators*	[10,20,30,40,50,100,150,200,300,500,600,800,1000]
	Minimum number of samples for an internal node	*min_samples_split*	[2, 3, 5, 8,9,12,15,20,40,50]
	Learning rate	*learning_rate*	[0.1,0.2,0.3,0.4,0.5,0.6,0.7,0.8,0.9]
	Maximum depth of individual regression estimators.	*max_depth*	[3,5,7,9,12,15,20,35,60,70,80]
	Number of features to consider for the best split	*max_features*	['auto', 'sqrt', 'log2']
Extreme Gradient Boosting Regressor	Number of boosted trees to fit	*n_estimators*	[10,20,30,40,50,60,70,80,90,100,200,300,500,600,700,800]
	Maximum tree depth for base learners	*max_depth*	[3,4,5,6,7,8,9,10,11,14,15,20,25,30,35,40,45,50,60,70,80]

Table 5(b). Optimum hyperparameters after grid search for each solar city, Australia (Blacktown, Central Victoria, Adelaide, and Townsville).

Model	Acronym	Blacktown	Central Victoria	Adelaide	Townsville
	max_depth	6	5	6	9
Decision Tree	*min_samples_split*	5	3	10	3
	max_features	auto	auto	auto	auto
	n_estimators	20	180	90	50
	max_depth	15	15	15	10
Random Forest Regressor	*min_samples_split*	3	3	3	5
	max_features	auto	auto	auto	auto
	n_estimators	300	1000	100	100
	min_samples_split	12	15	50	40
Gradient Boosting Regressor	*learning_rate*	0.1	0.1	0.1	0.1
	max_depth	3	3	9	9
	max_features	auto	auto	NA	auto
Extreme Gradient Boosting Regressor	*n_estimators*	200	200	200	80
	max_depth	3	3	3	4

Note: NA refers to a parameter that is not applicable.

3.4. Model Performance Criteria

To evaluate the performance of the proposed deep learning models against their comparative counterparts, statistical metrics were employed. Commonly-used metrics like *RMSE*, *MAE* and Pearson's correlation coefficient (*r*), together with skill score metrics ($RMSE_{ss}$) defined in Equation (7) are as follows.

$$RMSE_{ss} = 1 - \frac{RMSE_{\mathcal{M}}}{RMSE_{\mathcal{P}}} \tag{7}$$

where $\mathcal{M}$ refers to the error (*RMSE*) obtained in the predicted results employed to assess model performance. Here, $\mathcal{P}$ is the *RMSE* of a persistence model. The persistence model, which is also called the naive predictor, considers that the *GSR* at $t + 1$ equals the *GSR* at t. The interpretation of this metric is that a value of $RMSE_{ss}$ close to zero will indicate that the performance of the model is similar to that of the persistence model.

By contrast, if this metric is a positive value, the models under study are likely to outperform the persistence model (which is the baseline), whereas if $RMSE_{ss}$ attains a negative value, then the persistence model is likely to be better than the models under study. This study also utilized the normalized performance indicators based on the Nash–Sutcliffe coefficient (E_{NS}), Willmott's index (*WI*) and the Legate's and McCabe's index (*LM*), which provide advanced assessment of models relative to the E_{NS} and WI values.

In addition to these metrics, this study considered absolute percentage bias (*APB*) and Kling–Gupta efficiency (*KGE*) as key performance indicators for *GSR* prediction. The optimal value of *APB* is 0.0, with low-magnitude values indicating accurate model simulation; whereas, *KGE* is a model evaluation criterion that can be decomposed into the contribution of the mean, variance and correlation on model performance [91]. A *KGE* value of unity is considered as the perfect fit. Similarly, underestimation bias and overestimation bias of models are represented by positive and negative values of *APB*, respectively.

The mathematical derivation is as follows:

$$E_{NS} = 1 - \frac{\sum_{i=1}^{N} \left[GSR^m - GSR^p \right]^2}{\sum_{i=1}^{N} \left[GSR^m - \langle GSR^m \rangle \right]^2} \tag{8}$$

$$WI = 1 - \frac{\sum_{i=1}^{N} \left[GSR^m - GSR^p \right]^2}{\sum_{i=1}^{N} \left[|(GSR^p - \langle GSR^m \rangle)| + |(GSR^m - \langle GSR^m \rangle)| \right]^2} \tag{9}$$

$$LM = 1 - \frac{\sum_{i=1}^{N} |GSR^m - GSR^p|}{\sum_{i=1}^{N} |GSR^m - \langle GSR^m \rangle|} \tag{10}$$

$$APB = \left[\frac{\sum_{i=1}^{N} (GSR^m - GSR^p) * 100}{\sum_{i=1}^{N} GSR^m} \right] \tag{11}$$

$$KGE = 1 - \sqrt{(r-1)^2 + \left(\frac{<GSR^p>}{<GSR^m>} - 1 \right)^2 + \left(\frac{CV_p}{CV_m} \right)^2} \tag{12}$$

where

r = correlation coeffecient

CV = coefficient of variation

where GSR^p and GSR^m are the predicted (i.e., estimated) and measured values, respectively, and $\langle\ \rangle$ refers to the average value of the respective data in the tested set.

4. Results and Discussion

In this section, the results generated by the DBN and DNN algorithms within the testing phase are presented to ascertain the appropriateness of the two deep learning methods used for *GSR* prediction, tested at diverse sites including Australia's solar cities. These were also benchmarked against single hidden layer (*i.e.*, ANN) and ensemble models (random forest regression (RF), extreme gradient boosting regression (XGBoost), Gradient Boosting Machine (GBM) and decision trees (DT). It is noteworthy that the results are only presented for DNN_{10} and DBN_2, the two optimized models in accordance with Tables 4 and 5.

Figure 3 shows this for Adelaide, where the optimal model was selected on the basis of the lowest *RRMSE*. It should be noted that for this location, the optimized models (DBN_{10} and DNN_2) (where SGD was used as the back-propagation algorithm) with input selected by the PSO approach were seen to yield a better result. Similarly, for the ANN model, the inputs screened by the nondominated sorting genetic algorithm (NSGA) feature selection approach appeared to be better, and similarly for the DT, GBM, RF and XGB models, inputs screened by the sequential forward selection (SFR), SA, sequential backward selection (SBR), step and NSGA appeared to yield a relatively low *RRMSE* compared to other feature selection methods.

The model designation is as follows: DBN_{10} = Deep Belief Network 10, $DNN2_{SGD}$ = Deep Neural Network 2 with SGD as back-propagation, ANN = neural network, DT = decision tree, RF= random forest regression, GBM = gradient boosting machine and XGBR= extreme gradient boosting regression.

In this study, a feature selection (FS) and sensitivity analysis process utilizing a Lowry plot (Figure 2) were combined to deduce the best FS method for *GSR* prediction. Note that the nomenclature of any model is designated as FS-(model name), for example, for the Adelaide study site, the model names are designated as $PSO-DBN_{10}$, $PSO-DNN2_{SGD}$, PSO-ANN, PSO-DT, PSO-GBM, PSO-RF and PSO-XGBR. The number that appears after the respective model name, for example, DBN_{10}, is used to represent Deep Belief Network Model Number 10, as deduced from Table 3. Similarly, the subscript (SGD, AdaGrad, Adam) in the model names represent the back-propagation algorithm that was used in training the neural network model. For the ANN, however, only one back-propagation algorithm (LM), the most popular algorithm, was used, so the subscript is not mentioned.

Table 6 shows the training root mean squared error generated for different FS algorithms integrated with deep learning and its respective comparative algorithms, required to select the critical MODIS-derived predictors. In accordance with this evaluation, for the DBN_{10} model, the GA appeared to be the best FS algorithm for the study site Blacktown, whereas the relief algorithm was the best for the case of Central Victoria, and the PSO algorithm was the best for both the Adelaide and Townsville study sites. However, when the DNN_2 model was evaluated, root mean squared errors were slightly higher for each FS algorithm compared to those obtained from DBN_{10}, but these predictive errors remained much lower than all single hidden layer and ensemble models, thus confirming the superiority of deep learning over the less sophisticated models. When both deep learning models (*i.e.*, DBN_{10} and DNN_2) were evaluated individually, there appeared to be a clear consensus that the DBN_{10} model exceeded the performance of DNN_2, used with its best FS approach.

Table 7 compares deep learning models *vs*. the counterpart models in the testing phase, measured by the correlation coefficient (*r*), root mean squared error (*RMSE*), mean absolute error (*MAE)* and skill score (*$RMSE_{ss}$*). As mentioned earlier, only the optimally-trained models with the lowest *MAE* and *RMSE*, the highest *r* and the *$RMSE_{ss}$* values are shown. Between the deep learning and comparative (SHL and ensemble) models, the DBN model yielded better *GSR* predictions for all four solar cities. This is evident, for example, when comparing the DBN accuracy statistics (*i.e.*, *r* ≈ 0.994, *RMSE* ≈ 0.546 MJ·m^{-2}·day^{-1}, *MAE* ≈ 0.450 MJ·m^{-2}·day^{-1} and *$RMSE_{ss}$* ≈ 0.824 for Blacktown GA-

DBN$_{10}$) with the equivalent ANN and GBM models result statistics ($r \approx 0.989$, *RMSE* ≈ 0.739 MJ m^{-2}day^{-1}, M*AE* ≈ 0.536 MJ·m^{-2}·day^{-1} and RM*SE*$_{ss}$ ≈ 0.739 for Blacktown GA-ANN and $r \approx 0.988$, *RMSE* $\approx$ 0.664 MJ·m^{-2}·day^{-1}, *MAE* ≈ 0.568 0 MJ·m^{-2}·day^{-1} and *RMSE*$_{ss}$ ≈ 0.787 for Blacktown GA-GBM). Comparatively better results for DBN$_{10}$ models were also seen for all of the other solar cities, confirming the reliability of this deep learning approach, as a viable estimator *GSR*, with implications for long-term solar energy assessments.
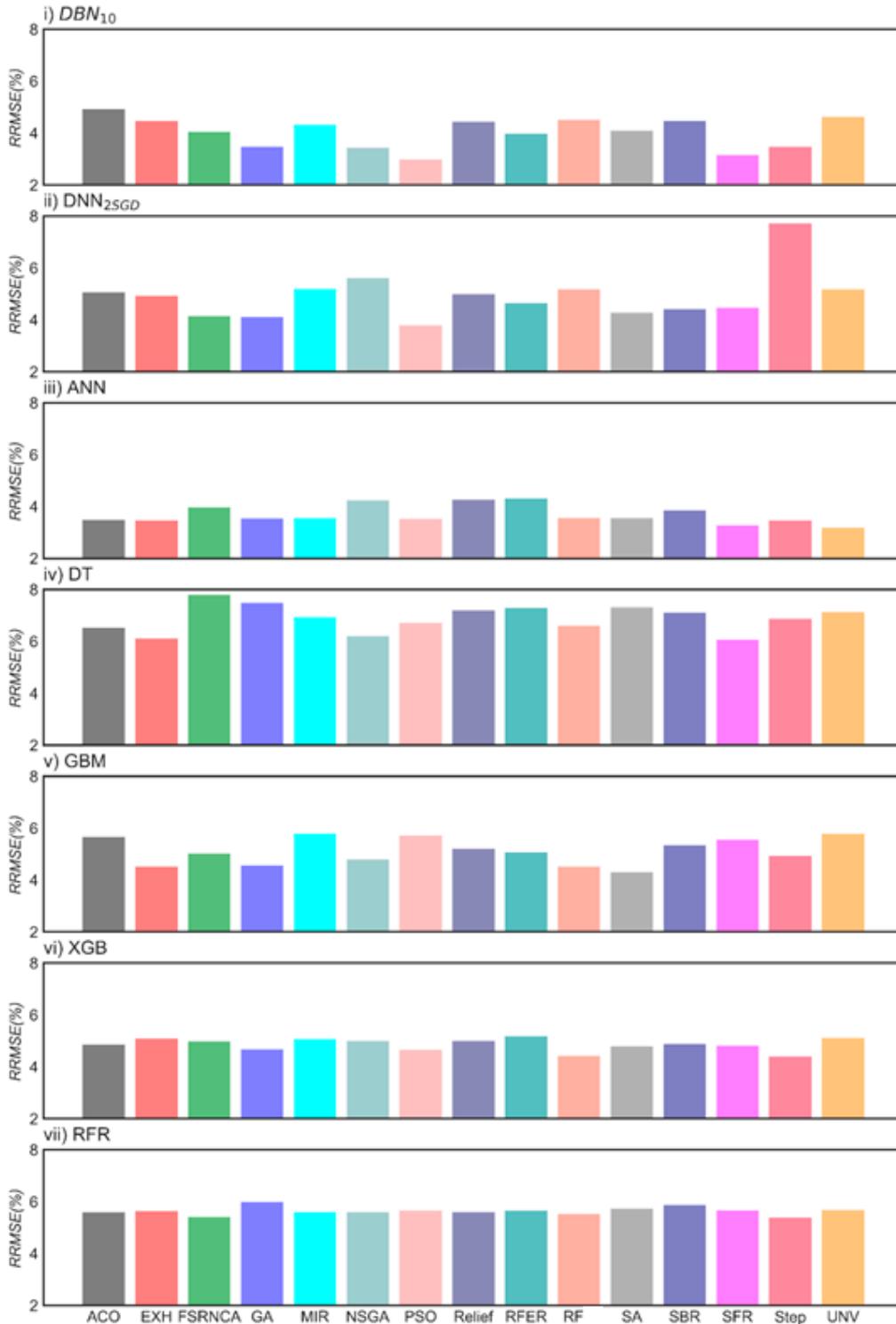
Figure 3. Relative root mean square error (*RRMSE %)* in the model's testing phase, illustrated for a selected solar city, Adelaide, to identify the accurate performance using

different feature selection algorithms. Note: For acronyms and model names, refer to Table 1, 2, 3 and 5.

Table 6. Evaluating the training performance of DBN$_{10}$ and DNN$_2$ *vs.* the counterpart models in terms of their best feature selection methods, as measured by the relative root mean squared error (*RRMSE*, %) used for long-term *GSR* predictions.

Predictive Model Type		Acronym	Approach for Feature Selection and Performance Error	Australia's Solar City			
				Blacktown	Central Victoria	Adelaide	Townsville
Deep Learning Models		**DBN$_{10}$**	Best Feature Selection Algorithm	**GA**	**Relief**	**PSO**	**PSO**
			RRMSE	**3.279**	**3.7133**	**2.988**	**3.791**
		DNN$_2$	Best Feature Selection Algorithm	**GA**	**Relief**	**PSO**	**EXH**
			RRMSE	**3.66**	**4.8296**	**3.774**	**3.899**
Single Hidden Layer ANN, Decision Tree and Ensembles Models		ANN	Best Feature Selection Algorithm	ACO	FSRNCA	NSGA	NSGA
			RRMSE	4.39	6.252	3.449	4.666
		DT	Best Feature Selection Algorithm	RF	FSRNCA	SFR	Step
			RRMSE	6.49	9.7043	6.055	6.036
		GBM	Best Feature Selection Algorithm	SFR	Relief	SA	RFER
			RRMSE	3.755	6.177	4.302	4.959
		RF	Best Feature Selection Algorithm	ACO	UVR	Step	RFER
			RRMSE	4.884	7.0976	5.375	5.65
		XGBR	Best Feature Selection Algorithm	NSGA	NSGA	Step	RFER
			RRMSE	4.2808	6.3964	4.392	4.644

Note: The best models (DNN$_2$ and DBN$_{10}$) are highlighted in blue and **boldfaced,** and the symbols are as per Table 1 and 2a.

Table 7. Comparison of deep learning models vs. counterpart models. The best model is highlighted in blue and **boldfaced (**DBN$_{10}$ represents the optimal model, to accord with Table 3, and DNN$_{2\,Nadam}$ represents the optimal model, to accord with Table 5, trained with the Nadam-type back-propagation algorithm.

Australia's Solar City	Model	*r*	*RMSE* (MJ·m^{-2}·day^{-1})	*MAE* (MJ·m^{-2}·day^{-1})	*RMSE$_{ss}$*
Blacktown	DBN$_{10}$	**0.994**	**0.546**	**0.45**	**0.824**
	DNN$_{2Nadam}$	0.99	0.706	0.503	0.773
	ANN	0.989	0.739	0.536	0.739
	DT	0.955	1.309	0.979	0.579
	RF	0.982	0.798	0.635	0.744
	GBM	0.988	0.664	0.568	0.787
	XGBR	0.985	0.727	0.589	0.766
Adelaide	DBN$_{10}$	**0.997**	**0.503**	**0.426**	**0.863**
	DNN$_{2SGD}$	0.996	0.636	0.546	0.826
	ANN	0.997	0.653	0.529	0.824
	DT	0.985	1.063	0.791	0.713
	RF	0.989	0.895	0.652	0.754
	GBM	0.988	0.906	0.72	0.758
	XGBR	0.992	0.737	0.577	0.801
Central Victoria	DBN$_{10}$	0.996	**0.614**	**0.498**	**0.836**
	DNN$_{2SGD}$	**0.994**	0.798	0.592	0.787
	ANN	0.984	1.276	0.995	0.682
	DT	0.961	1.696	1.217	0.553
	RF	0.984	1.094	0.854	0.714
	GBM	0.988	0.942	0.799	0.753
	XGBR	0.987	0.992	0.825	0.74

	DBN$_{10}$	**0.974**	**0.773**	**0.627**	**0.718**
	DNN$_{2RMSProp}$	0.967	0.868	0.646	0.682
	ANN	0.972	0.991	0.858	0.641
Townsville	DT	0.951	1.181	0.973	0.572
	RF	0.95	1.212	0.971	0.559
	GBM	0.947	1.254	1.006	0.539
	XGBR	0.953	1.205	0.959	0.56
	DBN$_{10}$	**0.990**	**0.609**	**0.500**	**0.810**
	DNN$_{2RMSProp}$	0.987	0.752	0.572	0.767
	ANN	0.986	0.915	0.730	0.722
Average of 5 Study Sites	DT	0.963	1.312	0.990	0.604
	RF	0.976	1.000	0.778	0.693
	GBM	0.978	0.942	0.773	0.709
	XGBR	0.979	0.915	0.738	0.717

In conjunction with the statistical score metrics, the relative prediction errors were used to show the alternative "goodness-of-fit" for the predicted in relation to the observed *GSR* data. Figure 4 shows the radar plots in the model's testing phase for DNN$_2$ and DBN$_{10}$ in terms of the *RRMSE* (%) and *RMAE* % values. Note that these percentage errors were also used as alternative metrics to enable the model comparison at geographically-diverse sites [92]. It can be seen that the DBN model yielded high precision (with the lowest *RRMSE* and *RMAE*) followed by the comparative (SHL and ensemble) models.

For the optimal DBN model, the *RRMSE* and *RMAE* were found to be 3.279/2.763%, 2.989/3.124%, 3.713/3.572% and 3.792/3.175% for the study site Blacktown (GA-DBN$_{10}$), Adelaide (PSO-DBN$_{10}$), Central Victoria (Relief-DBN$_{10}$), and Townsville (PSO-DBN$_{10}$), respectively. On the other hand, the other deep learning model lagged behind the accuracy of DBN with 4.240/2.970%, 3.774/3.825%, 4.830/3.781% and 4.256/3.175% for Blacktown (GA-DNN$_{2Nadam}$), Adelaide (PSO-DNN$_{2SGD}$), Central Victoria (Relief-DNN$_{2SGD}$) and Townsville (PSO-DNN$_{2RMSProp}$), showing relatively good performance, compared to the SHL and ensemble models.

The SHL and ensemble models' performances were lower than those of the DBN and DNN models, except for the study site Adelaide, where the PSO-ANN model (*RRMSE/ RMAE*) was lower (3.702/3.442%) than the DNN model (*RRMSE/RMAE* ≈ 3.774/3.825%). In accordance with these outcomes, the relative measures concurred on the suitability of deep learning for *GSR* prediction at all four solar cities selected across Australia.
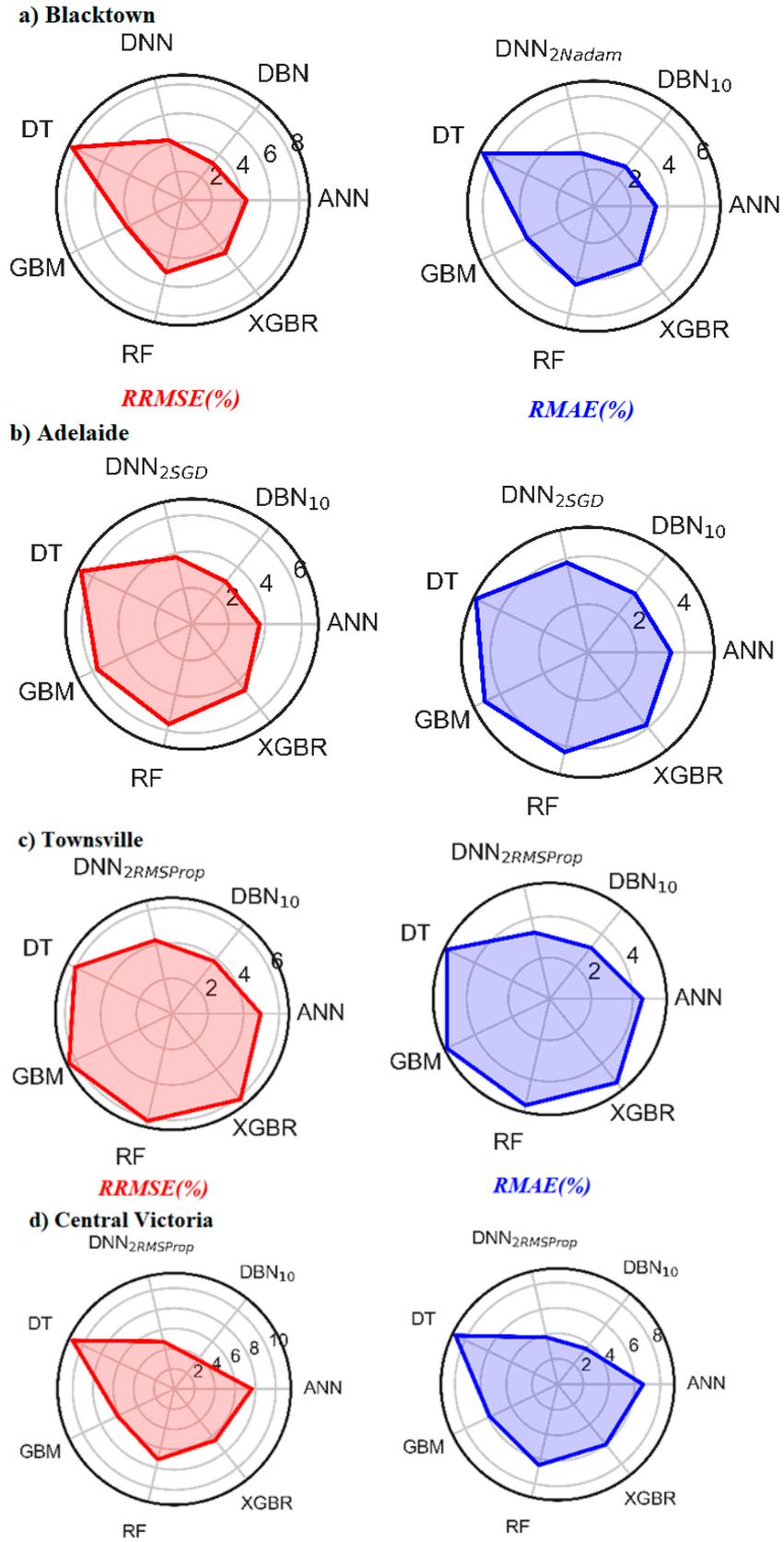
Figure 4. Radar plots in the model's testing phase for prediction of *GSR*, in terms of the relative root mean squared error (*RRMSE* %) and relative mean absolute error (*RMAE* %).

It is of interest to this study that a numerical quantification of model performance using Willmott's index *(WI)*, Nash–Sutcliffe *(E_{NS})* and Legates–McCabe's index *(LM)* was made where these metrics should ideally be unity for a perfect model. Subsequently, these results (Table 8) indicated that deep learning is able to attain a dramatic improvement in comparison to SHL and the ensemble model (Table 8).

The highest magnitude of *WI* ≈ 0.997, *E_{NS}* ≈ 0.995 and *LM* ≈ 0.933 was registered for the Blacktown study site for a GA-DBN$_{10}$ model. Intriguingly, the lowest value of *WI* ≈ 0.943, *E_{NS}* ≈ 0.891 and *LM* ≈ 0.689 was registered at the Townsville study site for the PSO-RF model. Further, the GA-DBN$_{10}$ model noted the increment in *WI*, *E_{NS}* and *LM* by 5.2%, 9.06% and 21.02%, respectively, compared to the counterpart (SHL and ensemble) models for the Blacktown study site. A similar trend was also demonstrated for the other solar cities. Hence, it is evident that a deep learning model has a better potential to predict *GSR* over long-term periods.

Table 8. Performance of deep learning models (*i.e.*, DNN$_{10}$, DBN$_{2Nadam}$) with respect to their comparative counterpart models. The best model is highlighted in bold.

Australia's Solar City	Model	Model Performance Metrics		
		WI	E_{NS}	LM
Blacktown	DBN$_{10}$	0.993	0.987	0.893
	DNN$_{2Nadam}$	0.989	0.979	0.88
	ANN	0.989	0.977	0.875
	DT	0.944	0.905	0.738
	RF	0.981	0.965	0.83
	GBM	0.986	0.976	0.848
	XGBR	0.984	0.971	0.843
Adelaide	DBN$_{10}$	0.997	0.995	0.933
	DNN$_{2SGD}$	0.996	0.991	0.915
	ANN	0.995	0.989	0.907
	DT	0.982	0.968	0.848
	RF	0.987	0.977	0.875
	GBM	0.987	0.977	0.862
	XGBR	0.991	0.985	0.889
Central Victoria	DBN$_{10}$	0.996	0.992	0.922
	DNN$_{2RMSProp}$	0.994	0.987	0.908
	ANN	0.984	0.964	0.837
	DT	0.958	0.923	0.773
	RF	0.982	0.968	0.84
	GBM	0.986	0.976	0.851
	XGBR	0.985	0.974	0.846
Townsville	DBN$_{10}$	0.975	0.949	0.786
	DNN$_{2RMSProp}$	0.969	0.936	0.78
	ANN	0.965	0.919	0.713
	DT	0.957	0.904	0.699
	RF	0.949	0.898	0.7
	GBM	0.943	0.891	0.689
	XGBR	0.948	0.9	0.703

Note: DBN$_{10}$ means the DBN model as per Table 3's configuration, and DNN$_{2Nadam}$ means the DNN model as per Table 5 with Nadam as the back-propagation algorithm.

In this paper, a comprehensive evaluation of the deep learning approach for *GSR* predictions was made in terms of the absolute percentage bias (*APB*, %) and Kling–Gupta efficiency (*KGE*) in the testing phase (Figure 5). The evaluation of *KGE* and *APB* for these solar cities showed that the DBN_{10} model constituted the best performing approach.

For example, $KGE \geq 0.99$ and $APB \leq 0.025$ for the case of Blacktown (GA-DBN_{10}), Adelaide (PSO-DBN_{10}), Central Victoria (Relief-DBN_{10}) and Townsville (PSO-DBN_{10}). Indeed, this plot shows that the magnitude of *KGE* was oriented toward unity, and the magnitude of *APB* was oriented toward zero for all deep learning models for all solar cities in consideration. Concurrent with the earlier findings, the deep learning model can be considered as a trustworthy and powerful tool for the prediction of long-term *GSR*, at least by the evidence generated so far.

Further insights were gained by checking the correspondence between the predicted and actual *GSR*. Comparing the prescribed deep learning models with an earlier study using wavelet support vector machine models (W-SVM) [14] applied in Australia, it became evident that the precision of the present model was relatively good for prediction of daily averaged monthly global solar radiation. In fact, in that study, their W-SVM model for the Townsville study produced a regression line equation $GSR_{pred} = \mathbf{0.849} \times GSR_{obs} + \mathbf{3.02}$, whereas the present deep learning model generated $GSR_{pred} = \mathbf{0.969} \times GSR_{obs} + \mathbf{0.678}$ and $GSR_{pre} = \mathbf{0.939} \times GSR_{obs} + \mathbf{1.196}$ for DBN (PSO-DBN_{10}) and DNN (PSO-$DNN2_{RMSProp}$), respectively. It is therefore clear that the prescribed approaches exceeded the performance of earlier studies.
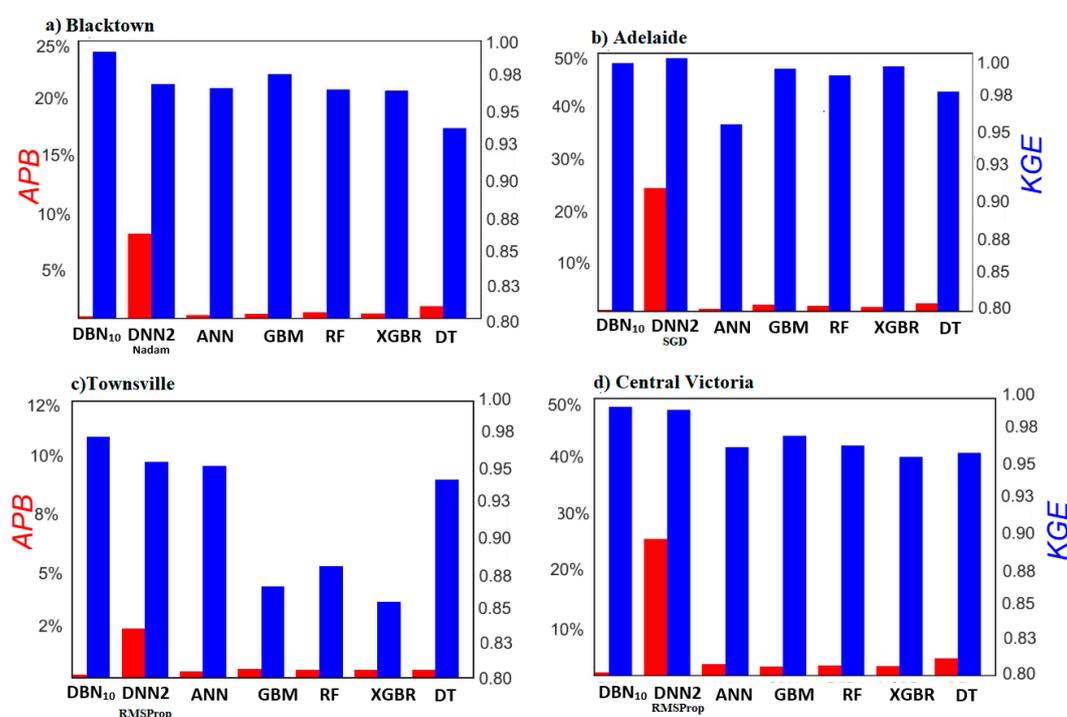
Figure 5. Bar chart showing a comparison of the optimal deep learning models (*i.e.*, DBN_{10} and $DNN2_{RMSProp}$) in terms of their absolute percentage bias (*APB*, %) and the Kling–Gupta efficiency (*KGE*) in the testing phase. *For notations and model names, please refer to Table 1, Table 2, Table 3, Table 5 and Table 7).*

To assess a model's stability for predicting *GSR*, the spread of the prediction errors is illustrated with the help of a violin plot [92] (Figure 6). In plotting this, all of the sites' actual and predicted *GSR* were considered. It should be noted that a violin plot is a synergistic combination of a box plot and density trace that is rotated and placed on each side to show the distribution shape of these data. The interquartile range is represented by a thick black bar in the centre, whereas 95% confidence intervals are represented by the thin black line or whisker, and the median is represented by the dot.

The shape of the violin displays the frequencies of the values. As can be seen on the figure (Figure 6) with a wider section of the plot, the prediction error (*PE*) generated by the DBN model had a high probability of a value of zero as compared to the benchmark models. Likewise, the median error (white dot) for a deep learning model was lower than that of comparative (SHL and ensemble) models, and the shape of distribution (extremely thin on each end and wide in the middle) indicates that the *PE* of DBN were highly concentrated around the median. Overall, it is noteworthy that the DBN model enjoyed superior performance relative to its comparative (SHL and ensemble) models tested for all four Australian solar cities.
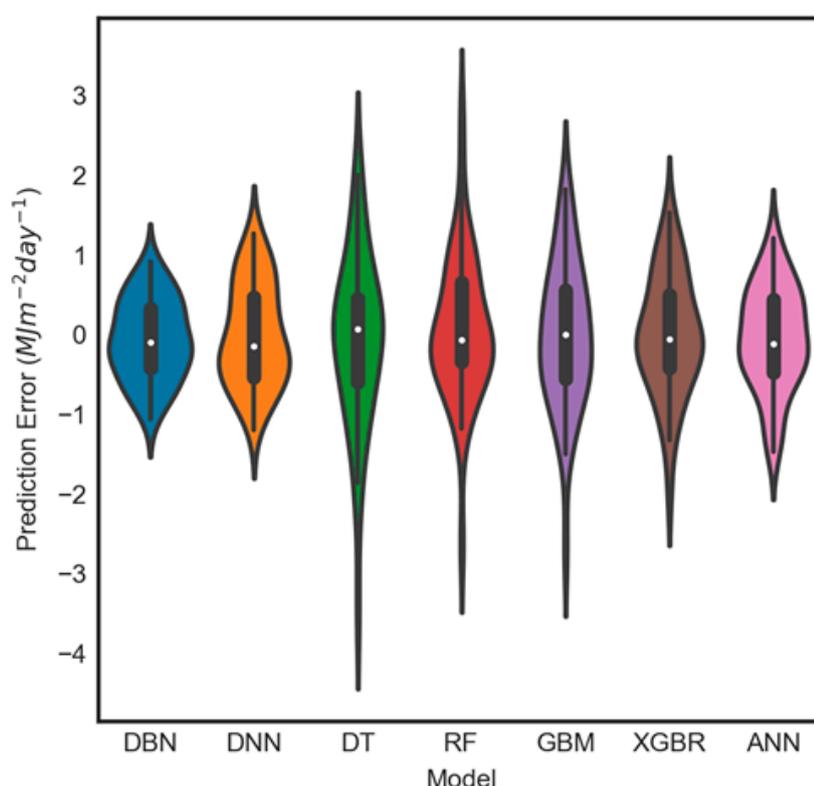
Figure 6. Violin plots of the prediction error (*PE*) generated by deep learning models (*i.e.*, DBN_{10} and $DNN2_{Nadam}$) compared with single hidden layer neuronal and decision tree-based models in the testing phase.

To draw a more conclusive argument on the suitability of the deep learning model for *GSR* prediction, Table 9 shows the prediction error (%), with its respective normalized frequency of the datum points, in each error bracket tested for each of the four solar cities. The normalized frequency is presented as a percentage of the predicted points in each error bracket in terms of the total data points in the testing period. Consistent with the earlier results, the most accurate prediction was obtained by using a deep learning model as the *PE* (%) attained the maximum value for the lowest range (*e.g.*, Townsville ≈ 81 % (DBN and DNN) within $[0 \leqslant |PE| < 4]$) compared to the comparative (SHL and ensemble) models (*e.g.*, Townsville ≈ 72.3 % (XGBR) within $[0 \leqslant |PE| < 4]$).

In Figure 7, the sensitiveness of clouds, water vapour and aerosols (*i.e.*, the three main contributors to the fluctuations in global solar radiation) including these GEM-SA critical parameters (Lowry plot, Figure 2) as the predictors of *GSR* are explored more closely for the case of the Adelaide study site. The four DBN models were tested with the cloud parameters (*i.e.*, *cfd*, *cfn*, *cotc*, *cotl*, *ctpm* and *cttd*), aerosol parameters (*asa* and *aod*), atmospheric water vapour (*awvl* and *awvm*) and GEM-SA critical parameters (*asa*, *cfd*, *awvl*, and *awvm*) to arrive at conclusive arguments. The model results were compared with the original model $PSO-DBN_{10}$ (Table 3).

From the graph (Figure 7), it is deduced that the PSO feature selection gave the best prediction with the lowest *RRMSE* and the highest values of *KGE*, *WI* and *r*. When only aerosol products from

the MODIS data repository were used as a potential input, the *RRMSE* appeared to increase by 65%, whereas the *WI*, *KGE* and *r*-values decreased by 66.8%, 67% and 66.9%, respectively. Similarly, with the cloud and water vapour product as a potential input, the *RRMSE* increased much greater than 600%, and *KGE*, *WI* and *r* decreased by more than 70%.

Furthermore, with only four critical parameters from GEM-SA (Lowry plot, *asa, cfd, awvl* and *awvm*) as an input, the *RRMSE* was lower than that of the aerosol, cloud and water vapour product as a potential input. Therefore, it is evident that the cloud and aerosol products were very important predictors and should not be neglected, for *GSR* predictions at the selected study sites.

Table 9. The percentage frequency of the absolute prediction errors, |PE|, in different error bands in the testing phase, encountered by the deep learning model within respect to its comparative counterpart models: a single hidden layer neural network (ANN) and ensemble models for Australia's solar cites. The *best model is highlighted in* blue/**bold**.

Prediction Error, |PE| (%)	Adelaide				Blacktown				Townsville				Central Victoria			
	DBN	**DNN**	ANN	XGBR	**DBN**	**DNN**	ANN	XGBR	**DBN**	**DNN**	ANN	XGBR	**DBN**	**DNN**	ANN	XGBR
0 ⩽ |PE| < 4	**86.40**	**88.60**	84.10	79.60	**88.60**	**84.10**	84.10	79.50	**81.00**	**81.00**	78.30	72.60	**81.80**	**79.50**	86.40	84.10
4 ⩽ |PE| < 5	11.40	4.60	4.60	11.40	2.30	11.40	4.60	11.40	7.10	14.30	12.20	12.20	11.40	11.40	2.30	11.40
5 ⩽ |PE| < 6	2.30	6.80	9.10	4.60	6.80	0.00	2.30	6.80	9.50	0.00	7.10	8.10	4.60	4.60	9.10	2.30
|PE| > 6	0.00	0.00	2.30	4.60	2.3	4.60	9.10	2.30	2.40	4.80	2.40	7.10	2.30	4.6	2.30	2.30
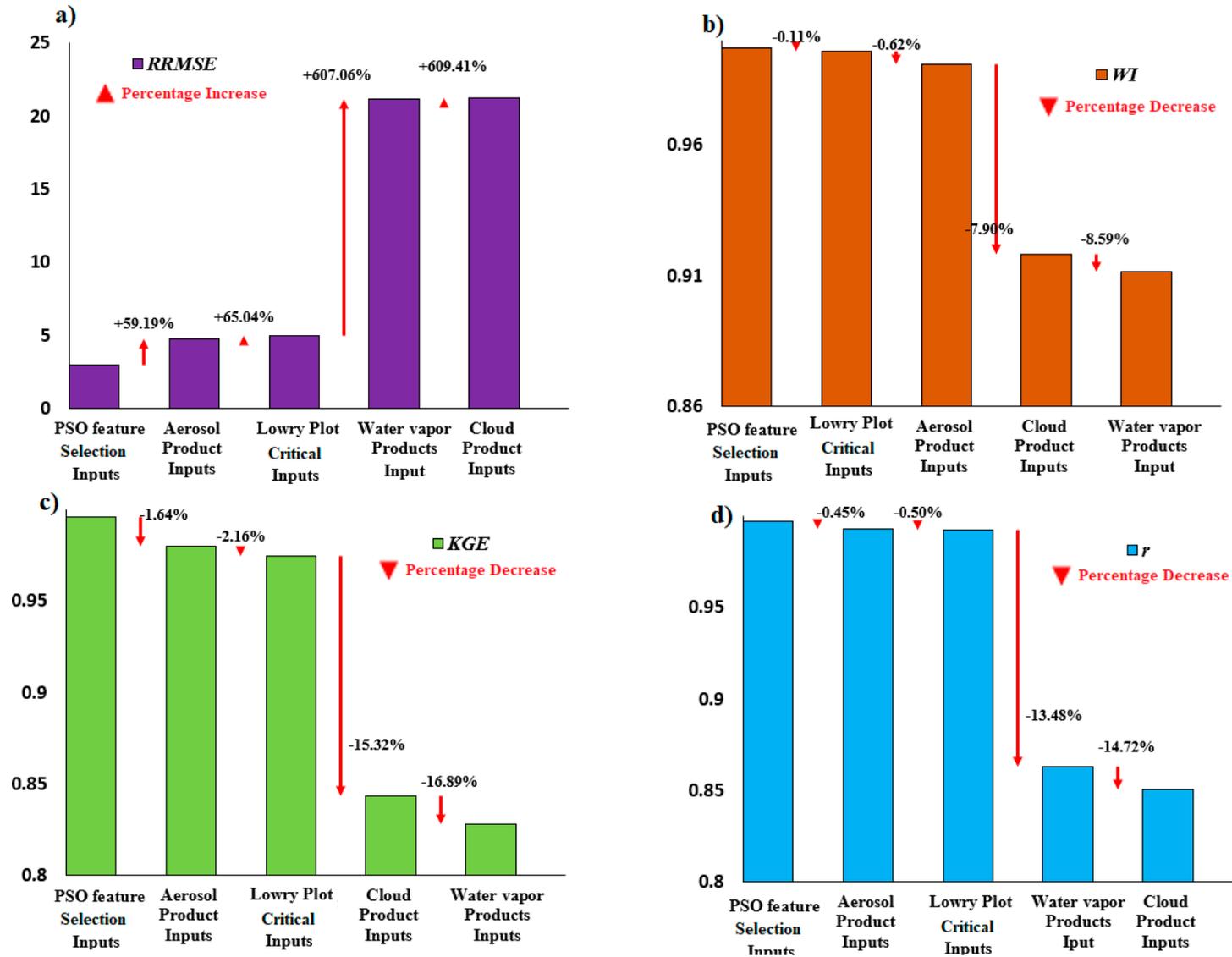
Figure 7. Sensitivity analysis of the relevant MODIS satellite-derived inputs variables: aerosol, cloud and water vapour, in terms of their: (**a**) relative root mean squared error (*RRMSE*), (**b**) Willmott's index (*WI*), (**c**) Kling–Gupta efficiency, (*KGE*), and (**d**) correlation coefficient (*r*).

5. Further Discussion

5.1. Comprehensive Evaluation of the Deep Learning Approach

It appears that DBN, DNN and also the ANN model (where the first two were designed with a deep learning approach, whereas the third model was designed with a single hidden layer neuronal system) have attained a better accuracy in comparison with all other data-driven models. As far as the error measurements were concerned, the *r*, however, was based on a linear relationship between observed and predicted *GSR* and, therefore, can be limited in its capacity to provide a robust model since it standardizes the observed and predicted means and variances [13]. However, *RMSE* and *MAE*, used in this study, can provide information about the predictive skill, whereas *RMSE* measures the goodness-of-fit relevant to high values, and *MAE* is not weighted towards high(er) magnitude or low(er) magnitude, but instead evaluates all deviations from observed, in both an equal manner and regardless of sign [9]. It is for this reason that all of these metrics were used to evaluate the deep learning models for GSR prediction.

It is important to note that while *RMSE* can assess the model with higher skill compared to the correlation coefficient, this metric is computed on squared differences [9]. Thus, performance assessment is biased in favour of the peaks and high(er) magnitude events that will in most cases exhibit the greatest error and be insensitive to low(er) magnitude sequences [14]. Consequently, the *RMSE* can be more sensitive than other performance metrics.

To overcome this issue, in this study, relative errors, root mean squared error (*RRMSE*) and mean absolute error (*RMAE*) (Figure 6) were utilized to describe the model over the range of statistically-different *GSR*, making it possible to compare the models evaluated for geographically- (and climatically-) diverse sites, where *MAE* and *RMSE* alone do not make sense. Notably, although the *r* (≥0.98) value was similar for DBN, DNN and ANN, in terms of *RRMSE* and *RMAE*, DBN and DNN outperformed all of the other models. Furthermore, *KGE* was quite high (≥97.5%) (Figure 7) for the DBN model for all locations as compared to DNN, ANN and the other models. Note that *KGE* gives more weight to the mismatch between observed and predicted *GSR* for high *GSR* because it squares the difference. Additionally, *APB* was low (≤3%) for DBN, measuring the tendency of predicted *GSR* to be larger or smaller than observations.

Comprehensively considering the prediction accuracy in terms of *RRMSE, RMAE, KGE* and *APB* of deep learning models, the DBN model can be of high utility for predicting long-term *GSR* using remote sensing data under different climatic conditions in Australia and, perhaps, elsewhere with similar climatic conditions.

5.2. Comparison with Related Research Work

In spite of good performance attained by deep learning approaches, as evidenced by statistical metrics and visual analysis, we further evaluated the models with respect to results from other studies. One such study was the work of [44] that validated a DBN model for daily *GSR* predictions in the Lhasa region in China using weather station fields (*i.e.*, wind speed, sunshine duration, air dry-bulb temperature, air relative humidity) for the period 1994–2015. In concurrence with the present study (Table 7), that study also concluded that the DBN approach constituted the best model as it generated a relatively low mean absolute bias error, *RMSE* and a high *r*-value (*i.e.*, 1.2709 MJ m^{-2} day^{-1}, 1.6765 MJ m^{-2} day^{-1} and 0.960).

Further comparison can be made in terms of another relevant study [93] where a DNN model was employed to estimate daily *GSR* over 30 stations in Turkey. The astronomical factor, extra-terrestrial radiation and climatic variables, sunshine duration, cloud cover, minimum temperature and maximum temperature, were used as input variables with data from 34 stations spanning from 2001–2007 used for training and testing the models. Their proposed DNN model yielded a high coefficient of determination (r^2 = 0.980) and low *RMSE* (0.780 MJ m^{-2} day^{-1}) and *MAE* (0.610 MJ m^{-2} day^{-1}), which stands within the range of the present study (Table 7).

The results from two latest research works were, however, the closest available comparison of deep learning models for *GSR* prediction. When directly comparing the prediction metrics, our DBN model outperformed by a noticeable margin, with a lower *RMSE* (≤ 0.503 MJ·m^{-2}·day^{-1}) and *MAE* (≤ 0.426 MJ·m^{-2}·day^{-1}) and high *r* (≥0.994) values (Table 7). Moreover, the two recent research projects used weather station data (ground-based), whereas the present study used freely-available remote sensing data as potential inputs (Table 2). Furthermore, a feature selection and sensitivity analysis of the MODIS predictors were not performed in the compared research works, whereas this study used fifteen feature selection algorithms (Table 2a) including the GEM-SA method for sensitivity analysis of the MODIS-derived predictors.

In terms of using deep learning models for *GSR* prediction in Australia, this study is the first of its kind to demonstrate the merits of this algorithm with respect to a neural network model used previously (*e.g.*, ANN). For example, a study by Deo and Sahin [13] that utilized MODIS land surface temperature within an ANN model for long-term *GSR* predictions and generated an average *RMSE* of 1.23 MJ·m^{-2} over a group of seven sites in regional Queensland had errors far exceeding a current lower average *RMSE* of 0.609 MJ·m^{-2}. Furthermore, that study had generated an average *MAE* of 1.02 MJ·m^{-2} in contrast to a lower error value of 0.50 MJ·m^{-2}.

The comparison shows that the proposed study can be considered as a significant advancement over earlier studies performed in Australia that have used satellite-derived variables, but did not apply deep learning. In this context, a deep learning approach may be adopted for long-term solar radiation modelling and future solar energy exploration.

5.3. Recommendation for Further Research

This study supports the significant merits of a deep learning predictive model to attain greater precision in predicting long-term *GSR*. Further, the study also provided a significant guideline for selecting appropriate models for *GSR* prediction in terms of the predictive accuracy in different climatic zones in Australia and also may be applicable elsewhere where similar climatic conditions prevail around the world. However, the scope of this study was restricted in terms of the prediction horizon, authenticating deep learning for monthly averaged, daily *GSR, i.e.*, long-term period.

6. Conclusion

Deep learning models were developed for Australia's solar cities (*i.e.*, Adelaide, Blacktown, Townsville, and Central Victoria) to estimate long-term *GSR*. These cities are heterogeneously distributed and represent a significant variation in their climatic conditions. In order to predict the monthly averaged daily *GSR* as an output, publicly-available MODIS satellite data (aerosol, cloud and water vapour) from GIOVANNI were extracted as the most relevant predictors. Fifteen different wrapper and filter-based feature selection algorithms were applied, with sensitivity analysis of all MODIS-derived predictors using GEM-SA to select the optimum input for the prediction of *GSR*. The data were segregated 80% for training, and 20% data were used for testing. A total of 180 deep belief networks (12 DBN models, 15 feature selections) and 630 deep neural network models were developed for each site. The developed models were benchmarked with single hidden layer and ensemble models including neural network, gradient boosting machine, extreme gradient boosting regression, decision tree and random forest regression models.

A holistic evaluation via statistical metrics and diagnostic plots revealed that the DBN model generated superior prediction in comparison with the benchmark models (viz., ANN, GBM, XGBR, DT and RF). The site comparison showed that the DBN model had the best performance at Blacktown (Table 7) with the lowest *RRMSE* ≈ 2.988% and *RMAE* ≈ 2.76% and highest *r* ≈ 0.994 and *RMSE*$_{ss}$ ≈ 0.824 in predicting *GSR*. Similarly, the DBN model outperformed all the benchmark models for all sites (Figure 5) in terms of absolute percentage bias and Kling–Gupta efficiency (e.g., *KGE* ≈ 0.992 and *APB* ≈ 0.027 for Blacktown using the DBN model, *KGE* ≈ 0.855 and *APB* ≈ 0.049 for Townsville using the XGBR model). Furthermore, the regression plot of actual versus predicted *GSR* demonstrated that, with a slope closer to unity and an intercept closer to zero, the DBN was

best in *GSR* estimation and even outperformed the previous study [93] using a W-SVM model for *GSR* estimation at Townsville. In addition to this, the sensitivity analysis of the predictor variables demonstrated that aerosol, cloud, and water vapour parameters as input parameters played a significant role in the prediction of *GSR* (Figure 7). This is a clearly understandable finding, as cloud and aerosol have obvious noticeable effects on sky brightness during daylight hours.

The findings of this study ascertained that with appropriate feature selection (such as PSO, GA and GEM-SA for sensitivity analysis), the deep learning model effectively captured the nonlinear dynamics and interactions amongst the input parameters and *GSR* in generating optimally-combined and -stabilized predictions for all four study sites. The DL model yielded good results for estimating monthly averaged daily *GSR*, either better than or comparable to many previous studies reported in the literature. One can conclude that the method derived here can be implemented as a suitable alternative and be successfully applied to similar regions.

Author Contributions: Conceptualization, S.G.; methodology, S.G.; software, S.G.; validation, S.G.; formal analysis, S.G.; resources, S.G.; writing—original draft preparation, S.G.; writing—review and editing. R.C.D., N.R., J.M; supervision, R.C.D.

Conflicts of Interest: The authors declare no conflict of interest.

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